# Relativistic particles and the geometry of 4-D null curves 

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#### Abstract

We study actions in $(d+1)$-dimensions associated with null curves, mainly when $d=3$, whose Lagrangian is a linear function on the curvature of the particle path, showing that null helices are always possible trajectories of the particles. We find Killing vector fields along critical curves of the action which correspond to the linear and the angular momenta of the particle. They provide two constants of the motion which can be interpreted in terms of the mass and the spin of the system. Moreover, we are able to integrate both the Euler-Lagrange equations and the Cartan equations in cylindrical coordinates around a certain plane.


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## 1. Introduction

In the past fifteen years, many interesting papers concerning Lagrangians describing spinning particles have been published (see e.g. [1-11] and references therein). In the general situation, as is well known, one has to provide the classical model with the extra bosonic variables. To this end, an interesting hypothesis deals with Lagrangians on higher geometrical invariants to supply those extra degrees of freedom. This approach has the interesting point of view that the spinning degrees of freedom are encoded in the geometry of its world trajectories. The Poincaré and invariance requirements imply that an admissible Lagrangian density $F$ must depend on the extrinsic curvatures of curves in the background gravitational field. In particular, the Lagrangians depending on the first and second curvatures have been intensively studied in the last twenty years. At the beginning those systems were studied as toy models of rigid strings and $(2+1)$-dimensional field theories with the Chern-Simons term, but shortly after, mainly due to the papers by Plyushchay, those systems are of independent interest.

The actions considered before are defined on nonisotropic curves (spacelike or timelike), but on $(d+1)$-spacetimes one can also consider actions defined on null (lightlike) curves. The studies of Lagrangians on these curves begin in the late nineties by considering the simplest geometrical particle model associated with null paths in four-dimensional

[^0]Minkowski spacetime, [12], where the action is proportional to the pseudo-arc of the particle. The authors obtain the equations of motion and show that they are particular examples of null helices. The same authors consider in [13] this geometrical particle model associated with null curves in $(2+1)$-dimensions. Our main results have been recently exploited to provide a Lagrangian description of the dynamics of geometric models for null curves (see [14]). It is worth pointing out that our latest results form part of a general programme whose seminal paper was [15] (see also [16,17]).

The next step deals with a more complicated three-dimensional system where the action is a linear function on the curvature of the curve. In [18] the authors show that its mass and spin spectra are defined by one-dimensional nonrelativistic mechanics with a cubic potential. Recently, in [19] we obtain, using geometrical methods, a complete description of the relativistic particle paths.

This paper concerns actions in $(d+1)$-dimensions whose Lagrangian is a linear function on the curvature of the particle path. The paper is organized as follows: in Section 2 we present the model, whose action $\mathcal{L}: \Lambda \rightarrow \mathbb{R}$ is given by $\mathcal{L}(\gamma)=\int_{\gamma}\left(\mu k_{1}+\lambda\right) \mathrm{d} \sigma$, where $\mu$ and $\lambda$ are constants and $k_{1}$ stands for the first curvature of the null curve. The equations of motion for these Lagrangians are completely given in $(d+1)$-background gravitational fields. In Section 3 we solve the motion equations and get the null worldlines of the relativistic particles in cylindrical coordinates around a certain hyperplane $\Pi$. To this end, we distinguish two cases according to $\Pi$ is either non-null (space-like or timelike) or null. In Section 4 we give a complete description of critical curves, by using cylindrical coordinates, of the Lagrangian $\mathcal{L}(\gamma)$. Finally, Section 5 is devoted to discussion and concluding remarks.

## 2. The model and the equations of motion

Let $\mathbb{L}^{n}$ be an $n$-dimensional Lorentz-Minkowski space with background gravitational field $\langle$,$\rangle and Levi-Civita$ connection $\nabla$. First of all, we will describe the geometry of null curves in $\mathbb{L}^{n}$ in terms of the Cartan frame of the curve (see [20] for details).

Let $\gamma:[a, b] \rightarrow \mathbb{L}^{n}$ be a parametrized null Cartan curve such that the frame $\left\{\gamma^{\prime}(\sigma), \gamma^{\prime \prime}(\sigma), \ldots, \gamma^{(n)}(\sigma)\right\}$ is positively oriented, for all $\sigma \in[a, b], \sigma$ being the pseudo-arc parameter. Let us consider its corresponding Cartan frame $\left\{L=\gamma^{\prime}, W_{1}, N, W_{2}, \ldots, W_{n-2}\right\}$, where

$$
\begin{aligned}
& \langle L, L\rangle=\langle N, N\rangle=0, \quad\langle L, N\rangle=-1 \\
& \left\langle W_{i}, L\right\rangle=\left\langle W_{i}, N\right\rangle=0, \quad\left\langle W_{i}, W_{j}\right\rangle=\delta_{i j}
\end{aligned}
$$

The Cartan equations read

$$
\begin{align*}
& L^{\prime}=W_{1} \\
& W_{1}^{\prime}=-k_{1} L+N, \\
& N^{\prime}=-k_{1} W_{1}+k_{2} W_{2}, \\
& W_{2}^{\prime}=k_{2} L+k_{3} W_{3},  \tag{1}\\
& W_{i}^{\prime}=-k_{i} W_{i-1}+k_{i+1} W_{i+1} \quad i \in\{3, \ldots, n-3\}, \\
& W_{n-2}^{\prime}=-k_{n-2} W_{n-3},
\end{align*}
$$

where ()' means covariant derivative and $k_{i}$ are the Cartan curvatures of the curve. The fundamental theorem for null curves tells us that $\left\{k_{1}, \ldots, k_{n-2}\right\}$ determines completely the null curve up to Lorentzian transformations (see [20]). Even more, given functions $\left\{k_{1}, \ldots, k_{n-2}\right\}$, we can always construct a null curve, parametrized by the pseudo-arc length parameter $\sigma$, whose curvatures functions are precisely $\left\{k_{1}, \ldots, k_{n-2}\right\}$ (see [20, Theorem 3]). Then any local geometrical scalar defined along null curves can always be expressed as a function of its curvatures and derivatives.

In this section we analyse mechanical systems with Lagrangians depending linearly on the first curvature $k_{1}$ of the null curve. The space of elementary fields in this model is the set $\Lambda$ of all null Cartan curves satisfying given first order boundary data to drop out the boundary terms which appear when computing the first order variation of the action.

For the sake of simplicity $\gamma$ will also denote a variation of $\gamma$ by null curves $\gamma=\gamma(s, \omega):[a, b] \times(-\varepsilon, \varepsilon) \rightarrow \mathbb{L}^{n}$, where $\gamma(s, 0)$ is the reparametrization of $\gamma(\sigma)$. Associated with such a variation is the variational vector field $V(s)=$ $V(s, 0)$, where $V=V(s, \omega)=\frac{\partial \gamma}{\partial \omega}(s, \omega)$. Let $\eta$ be the differentiable function satisfying $\frac{\partial \gamma}{\partial s}(s, \omega)=\eta(s, \omega) L(s, \omega)$. Then we will write down $\gamma(\sigma, \omega), k(\sigma, \omega), V(\sigma, \omega)$, etc., for the corresponding pseudo-arc length parameter.

The actions $\mathcal{L}$ for the curve depend locally on its geometry and they possess various symmetries, both local and global. The local symmetry is reparametrization invariance and it restricts severely the form of $\mathcal{L}$. We consider the action $\mathcal{L}: \Lambda \rightarrow \mathbb{R}$ given by

$$
\begin{equation*}
\mathcal{L}(\gamma)=\int_{\gamma}\left(\mu k_{1}+\lambda\right) \mathrm{d} \sigma \tag{2}
\end{equation*}
$$

$\mu$ and $\lambda$ both being constant. The simplest action describing the motion of a particle is achieved when it is proportional to the pseudo-arc length parameter (i.e. $\mu=0$ ), which has been studied by Nersessian and Ramos in [12,13] when $n=2,3$. When the action is linear on the curvature of the particle path, some advances have been achieved in [18, 19].

A null curve $\gamma$ is said to be a critical point of the action $\mathcal{L}$ when

$$
\left.\frac{\mathrm{d}}{\mathrm{~d} \omega}\right|_{\omega=0} \mathcal{L}\left(\gamma_{\omega}\right)=\left.\frac{\mathrm{d}}{\mathrm{~d} \omega}\right|_{\omega=0} \int_{\gamma_{\omega}}\left(\mu k_{1}+\lambda\right) \mathrm{d} \sigma=0,
$$

for all variation throughout null curves $\gamma_{\omega}$ of $\gamma$.
Now we present a useful technical result.
Lemma 1. Using the above notation, the following assertions hold:
(a) $0=\left\langle\nabla_{L} V, L\right\rangle$;
(b) $\frac{\partial \eta}{\partial \omega}=V(\eta)=-\frac{1}{2} h \eta, \quad h=-\left\langle\nabla_{L}^{2} V, W_{1}\right\rangle ;$
(c) $\frac{\partial k_{1}}{\partial \omega}=\left\langle\nabla_{L}^{3} V, N\right\rangle+k_{1}\left\langle\nabla_{L} V, N\right\rangle+k_{1} h-\frac{1}{2} L(L(h))$;
(d) $\frac{\partial k_{2}}{\partial \omega}=\left\langle\nabla_{L}^{4} V, W_{2}\right\rangle+2 k_{1}\left\langle\nabla_{L}^{2} V, W_{2}\right\rangle+k_{1}^{\prime}\left\langle\nabla_{L} V, W_{2}\right\rangle+2 k_{2} h$.

A variational vector field $V$ along $\gamma$ which infinitesimally preserves the causal character, the pseudo-arc length parameter and the curvatures of $\gamma$ is said to be a Killing vector field along $\gamma$. Hence Killing vector fields along $\gamma$ are characterized by the following equations:

$$
\begin{aligned}
& \left\langle\nabla_{L} V, L\right\rangle=0, \\
& V(\eta)=0, \\
& V\left(k_{i}\right)=0, \quad i=1, \ldots, n-2 .
\end{aligned}
$$

As we will see, Killing vector fields will play an important role to integrate the Euler-Lagrange and Cartan equations. In particular, we are specially interested in the four-dimensional situation.
Proposition 2. Let $\gamma$ be an immersed null curve in $\mathbb{L}^{4}$. A vector field $V$ is Killing along $\gamma$ if and only if it extends to a Killing vector field $\tilde{V}$ on $\mathbb{L}^{4}$.

The same conclusion holds by considering a complete, simply connected, Lorentzian space form.
To compute the first-order variation of this action along the elementary fields space $\Lambda$, and so the field equations describing the dynamics of the particle, we use a standard argument involving some integrations by parts. Then the Cartan equations yield

$$
\begin{equation*}
\mathcal{L}^{\prime}(0)=\frac{1}{2}[\Omega(\gamma, V)]_{a}^{b}-\frac{1}{2} \int_{a}^{b}\left\langle V, \mathcal{E}_{1}(\gamma) L+\mathcal{E}_{2}(\gamma) W_{2}+\mathcal{E}_{3}(\gamma) W_{3}+\mathcal{E}_{4}(\gamma) W_{4}\right\rangle \mathrm{d} \sigma, \tag{3}
\end{equation*}
$$

where

$$
\begin{align*}
& \mathcal{E}_{1}(\gamma)=\mu k_{1}^{\prime \prime \prime}+2 \mu k_{2} k_{2}^{\prime}+3 \mu k_{1} k_{1}^{\prime}-\lambda k_{1}^{\prime} \\
& \mathcal{E}_{2}(\gamma)=2 \mu k_{2}^{\prime \prime}-k_{2}\left(2 \mu k_{3}-\mu k_{1}+\lambda\right), \\
& \mathcal{E}_{3}(\gamma)=2 \mu\left(k_{2} k_{3}^{\prime}-2 k_{2}^{\prime} k_{3}\right),  \tag{4}\\
& \mathcal{E}_{4}(\gamma)=k_{2} k_{3} k_{4},
\end{align*}
$$

and the boundary term reads

$$
\begin{equation*}
\Omega(\gamma, V)=\left\langle\nabla_{L}^{3} V, \mu W_{1}\right\rangle+\left\langle\nabla_{L}^{2} V,-\mu k_{1} L+3 \mu N\right\rangle+\left\langle\nabla_{L} V,\left(\mu k_{1}+\lambda\right) W_{1}-\mu k_{2} W_{2}\right\rangle+\left\langle V, P_{1}\right\rangle, \tag{5}
\end{equation*}
$$

where $P_{1}$ is the vector field given by

$$
\begin{equation*}
P_{1}=\left(\mu k_{1}^{\prime \prime}+\mu k_{1}^{2}-\lambda k_{1}\right) L-\mu k_{1}^{\prime} W_{1}+\left(\mu k_{1}-\lambda\right) N+2 \mu k_{2}^{\prime} W_{2}+2 \mu k_{2} k_{3} W_{3}, \tag{6}
\end{equation*}
$$

and $V$ stands for a generic variational vector field along $\gamma$.
To drop $[\Omega(\gamma, V)]_{a}^{b}$ we have to consider curves with the same endpoints and having the same Cartan frame there. Under these conditions, the first-order variation reads

$$
\mathcal{L}^{\prime}(0)=-\frac{1}{2} \int_{a}^{b}\left\langle V, \mathcal{E}_{1}(\gamma) L+\mathcal{E}_{2}(\gamma) W_{2}+\mathcal{E}_{3}(\gamma) W_{3}+\mathcal{E}_{4}(\gamma) W_{4}\right\rangle \mathrm{d} \sigma .
$$

As a consequence we have
Theorem 3. A null curve $\gamma \in \Lambda$ is critical for the linear action $\mathcal{L}(\gamma)$ in $\mathbb{L}^{n}$ if and only if the following statements hold:
(i) $W_{i}, N$ and $k_{j}$ are well defined along the whole trajectory;
(ii) The following differential equations are fulfilled:

$$
\mathcal{E}_{1}(\gamma)=0, \quad \mathcal{E}_{2}(\gamma)=0, \quad \mathcal{E}_{3}(\gamma)=0, \quad \mathcal{E}_{4}(\gamma)=0 .
$$

These equations are called the Euler-Lagrange equations. The following is an easy consequence from the last equation of (4):

Corollary 4. The critical points for the linear action $\mathcal{L}(\gamma)$ in $\mathbb{L}^{n}$ lie in a Lorentzian subspace of dimension not greater than five.

By considering the special case where the action is constant $(\mu=0)$, the Euler-Lagrange equations are reduced to

$$
-\lambda k_{1}^{\prime}=0, \quad-\lambda k_{2}=0, \quad k_{2} k_{3} k_{4}=0
$$

As a consequence we have
Theorem 5. The critical points for the constant action in $\mathbb{L}^{n}$ are just null helices in 3-dimensional Lorentzian linear subspaces.

We have made a more general treatment of Lagrangian in the 3-dimensional case (see [21]). There we have explicitly obtained all solutions for a linear action as well as got remarkable progress regarding other more difficult Lagrangians. Therefore, it seems reasonable to investigate the critical points of the linear action in the 4-dimensional case.

## 3. Linear action in $\mathbb{L}^{4}$

By reconsidering the action (2) in $\mathbb{L}^{4}$ we get $k_{3}=0$ (and $k_{4}=0$ ). Without loss of generality we normalize the constant $\mu$ by one and replace $k_{1}$ and $k_{2}$ by $k$ and $\tau$, respectively. Then the action (2) rewrite as

$$
\begin{equation*}
\mathcal{L}(\gamma)=\int_{\gamma}(k+\lambda) \mathrm{d} \sigma, \tag{7}
\end{equation*}
$$

and the Euler-Lagrange equations as

$$
\begin{equation*}
k^{\prime \prime \prime}+2 \tau \tau^{\prime}+(3 k-\lambda) k^{\prime}=0, \quad 2 \tau^{\prime \prime}+(k-\lambda) \tau=0 . \tag{8}
\end{equation*}
$$

Since the first equation can be easily integrated, the Euler-Lagrange equations state as

$$
\begin{equation*}
k^{\prime \prime}+\tau^{2}+\frac{3}{2} k^{2}-\lambda k+c=0, \quad 2 \tau^{\prime \prime}+(k-\lambda) \tau=0 \tag{9}
\end{equation*}
$$

where $c$ is an integration constant.
Now we are searching for nontrivial $(\tau \neq 0)$ solutions. We first find easy solutions when $k=\lambda$. Then $\tau$ is a nonzero constant and we get Cartan helices in $\mathbb{L}^{4}$. From now on, we will assume that $k \neq \lambda$. A straightforward computation shows that $P_{1}=\left(k^{\prime \prime}+k^{2}-\lambda k\right) L-k^{\prime} W_{1}+(k-\lambda) N+2 \tau^{\prime} W_{2}$ satisfies that $\nabla_{L} P_{1}=\mathcal{E}_{1}(\gamma) L+\mathcal{E}_{2}(\gamma) W_{2}$. Therefore, $P_{1}$ is a constant Killing vector field if and only if $\gamma$ is a solution of the Euler-Lagrange equations. We can apply the Noether argument relating rotational symmetries of $\mathcal{L}$ to constant of motions along $\gamma$. The variational vector field associated to variations generated by a one-parameter family of rotations is given by $\gamma \wedge Z_{1} \wedge Z_{2}$, where $Z_{1}$ and $Z_{2}$ are constant vector fields. As $\gamma$ satisfies the Euler-Lagrange equations and the action is rotationally invariant, we find

$$
\Omega\left(\gamma, \gamma \wedge Z_{1} \wedge Z_{2}\right)=\left\langle(k+\lambda) L \wedge W_{1} \wedge Z_{1}-2 \tau L \wedge W_{2} \wedge Z_{1}+2 W_{1} \wedge N \wedge Z_{1}+\gamma \wedge P_{1} \wedge Z_{1}, Z_{2}\right\rangle
$$

is constant for any constant vector field $Z_{2}$. Then the new vector field

$$
\begin{equation*}
Y=(k+\lambda) L \wedge W_{1} \wedge Z_{1}-2 \tau L \wedge W_{2} \wedge Z_{1}+2 W_{1} \wedge N \wedge Z_{1}+\gamma \wedge P_{1} \wedge Z_{1} \tag{10}
\end{equation*}
$$

is also constant along $\gamma$ for any constant vector field $Z_{1}$. In particular, we can replace $Z_{1}$ by $P_{1}$ to get the constant Killing vector field

$$
P_{2}=\left(2(k+\lambda) \tau^{\prime}-2 \tau k^{\prime}\right) L+2(k-\lambda) \tau W_{1}+4 \tau^{\prime} N+\left(2 k^{\prime \prime}+3 k^{2}-2 \lambda k-\lambda^{2}\right) W_{2}
$$

Observe that $P_{1}$ is orthogonal to $P_{2}$ unless both are null. This exceptional case will be considered later, where we will find explicit solutions. Then we will assume that one of them is non-null. By replacing $Z_{1}$ by $P_{2}$ into (10) we obtain a new constant vector field

$$
\begin{aligned}
X= & \left(2(k+\lambda) k^{\prime \prime}+4(k-\lambda) \tau^{2}+3 k^{3}+\lambda k^{2}-3 \lambda^{2} k-\lambda^{3}\right) L+8 \tau \tau^{\prime} W_{1} \\
& +\left(4 k^{\prime \prime}+6 k^{2}-4 \lambda k-2 \lambda^{2}\right) N+\left(8(k+\lambda) \tau^{\prime}-4 \tau k^{\prime}\right) W_{2}+P_{1} \wedge P_{2} \wedge \gamma .
\end{aligned}
$$

Set $J=X-P_{1} \wedge P_{2} \wedge \gamma$. As $J$ is the sum of a translational and a rotational vector fields, then $J$ is a Killing vector field. It is easy to check that $\left\langle P_{1}, P_{2}\right\rangle=\left\langle P_{2}, J\right\rangle=0$ and $\left\langle P_{2}, P_{2}\right\rangle=-\left\langle P_{1}, J\right\rangle$. We write down

$$
\left\langle P_{1}, P_{1}\right\rangle=\varepsilon_{1} p_{1}^{2}, \quad\left\langle P_{1}, J\right\rangle=\omega, \quad\left\langle P_{2}, P_{2}\right\rangle=\varepsilon_{2} p_{2}^{2},
$$

where $\varepsilon_{1}, \varepsilon_{2} \in\{-1,0,1\}, p_{1}=\left\|P_{1}\right\|, p_{2}=\left\|P_{2}\right\|$ and $\omega=-\varepsilon_{2} p_{2}^{2}$. The constants $p_{1}$ and $\omega$ represent the constants of motion of the relativistic particle, which correspond to the mass and spin of the system.

Proposition 6. Let $\gamma$ be a null curve in $\mathbb{L}^{4}$ with curvatures $k$ and $\tau$. Assume that $\left\{L, W_{1}, P_{1}, J\right\}$ is a set of linearly independent vector fields. Then $\gamma$ is a solution of the Euler-Lagrange equations if and only if $\left\langle P_{1}, P_{1}\right\rangle$ and $\left\langle P_{1}, J\right\rangle$ are both constant.

Proof. We first observe that $\gamma$ is a solution of the Euler-Lagrange equations if and only if the vector field $P_{1}$ along $\gamma$ satisfies $\nabla_{L} P_{1}=0$. To see that holds we have to show that $\left\langle\nabla_{L} P_{1}, P_{1}\right\rangle=\left\langle\nabla_{L} P_{1}, J\right\rangle=0$, because $\left\{L, W_{1}, P_{1}, J\right\}$ is a linearly independent system. A straightforward computation gives $\left\langle\nabla_{L} P_{1}, J\right\rangle=\left\langle P_{1}, \nabla_{L} J\right\rangle$, which leads to

$$
\frac{\mathrm{d}}{\mathrm{~d} \sigma}\left\langle P_{1}, P_{1}\right\rangle=2\left\langle\nabla_{L} P_{1}, P_{1}\right\rangle, \quad \frac{\mathrm{d}}{\mathrm{~d} \sigma}\left\langle P_{1}, J\right\rangle=2\left\langle\nabla_{L} P_{1}, J\right\rangle,
$$

and the result follows.
Therefore, asking for $\left\langle P_{1} \wedge J \wedge L, W_{1}\right\rangle \neq 0$, or said otherwise

$$
4(\lambda-k) \tau k^{\prime}-4 \tau^{\prime}\left(2 k^{\prime \prime}+k^{2}-2 \lambda k+\lambda^{2}\right) \neq 0
$$

the Euler-Lagrange equations are equivalent to those given by the constants of motion

$$
\begin{align*}
& 4\left(\tau^{\prime}\right)^{2}+\left(k^{\prime}\right)^{2}-2(k-\lambda)\left(k^{\prime \prime}+k^{2}-\lambda k\right)-\varepsilon_{1} p_{1}^{2}=0 \\
& 16\left((k+\lambda) \tau^{\prime}-\tau k^{\prime}\right) \tau^{\prime}-\left(2 k^{\prime \prime}+3 k^{2}-2 \lambda k-\lambda^{2}\right)^{2}-4(k-\lambda)^{2} \tau^{2}-\omega=0 \tag{11}
\end{align*}
$$

Combining (9) with (11) we get the equations of the motion

$$
\begin{align*}
& 4\left(\tau^{\prime}\right)^{2}+\left(k^{\prime}\right)^{2}+(k-\lambda)\left(k^{2}+2\left(\tau^{2}+c\right)\right)-\varepsilon_{1} p_{1}^{2}=0 \\
& 16\left((k+\lambda) \tau^{\prime}-\tau k^{\prime}\right) \tau^{\prime}-\left(2 \tau^{2}+2 c+\lambda^{2}\right)^{2}-4(k-\lambda)^{2} \tau^{2}-\omega=0 \tag{12}
\end{align*}
$$

To study this system we adopt a Hamiltonian point of view, because it reminds us a Hénon-Heiles system. To this end, we will provisionally introduce a more conventional notation. First, we set

$$
\begin{equation*}
q_{1}=k, \quad q_{2}=2 \tau, \quad p_{1}=k^{\prime}, \quad p_{2}=2 \tau^{\prime} . \tag{13}
\end{equation*}
$$

Then, the Euler-Lagrange equations (9) take the form

$$
\begin{equation*}
q_{1}^{\prime \prime}+\frac{1}{4} q_{2}^{2}+\frac{3}{2} q_{1}^{2}-\lambda q_{1}+c=0, \quad q_{2}^{\prime \prime}+\frac{1}{2} q_{2}\left(q_{1}-\lambda\right)=0 . \tag{14}
\end{equation*}
$$

The Hamiltonian function of the system is going to be the first constant of motion

$$
\begin{equation*}
H(q, p)=\frac{1}{2}\left\langle P_{1}, P_{1}\right\rangle=\frac{1}{2}\left(p_{1}^{2}+p_{2}^{2}-\lambda q_{1}^{2}-\frac{\lambda}{2} q_{2}^{2}\right)+\frac{1}{4} q_{1} q_{2}^{2}+\frac{1}{2} q_{1}^{3}+c\left(q_{1}-\lambda\right) . \tag{15}
\end{equation*}
$$

According to Eq. (14) we obtain

$$
\begin{aligned}
& p_{1}^{\prime}=q_{1}^{\prime \prime}=-\frac{1}{4} q_{2}^{2}-\frac{3}{2} q_{1}^{2}+\lambda q_{1}-c=-\frac{\partial H}{\partial q_{1}}, \\
& p_{2}^{\prime}=q_{2}^{\prime \prime}=-\frac{1}{2} q_{2}\left(q_{1}-\lambda\right)=-\frac{\partial H}{\partial q_{2}} .
\end{aligned}
$$

Thus, we conclude that the dynamics of this system obeys a classical two freedom degree Hamiltonian, the sum of kinetic and potential energies, where the potential is a cubic polynomial in the position variables $q_{1}, q_{2}$, that is,

$$
\begin{aligned}
& H=\frac{1}{2}\left(p_{1}^{2}+p_{2}^{2}\right)+V\left(q_{1}, q_{2}\right) \\
& V\left(q_{1}, q_{2}\right)=\frac{1}{2}\left(-\lambda q_{1}^{2}-\frac{\lambda}{2} q_{2}^{2}\right)+\frac{1}{4} q_{1} q_{2}^{2}+\frac{1}{2} q_{1}^{3}+c\left(q_{1}-\lambda\right) .
\end{aligned}
$$

Observe that when $c=0$ this system corresponds to a particular case of the general Hénon-Heiles Hamiltonian

$$
H=\frac{1}{2}\left(p_{1}^{2}+p_{2}^{2}+A q_{1}^{2}+B q_{2}^{2}\right)+D q_{1} q_{2}^{2}-\frac{C}{3} q_{1}^{3}
$$

where $A, B, C$ and $D$ are constant coefficients. The integrability of the system depends on the values of the coefficients. The equations of motion are integrable in three cases: (i) $A=B$ and $D=-C$; (ii) $6 D=-C$; and (iii) $16 D=-C$ and $6 A=B$. Note that Hamiltonian (15), when $c=0$, fit in case (ii), since $D=1 / 4$ and $C=-3 / 2$. By integrability here we mean existence of a second (global) integral of motion and, in this case, the Liouville theorem implies that the problem can be solved by quadratures. For a recent treatment of more general integrable cases we refer the reader to [22].

Going back to the primitive notation, we proceed to drop the curvature $\tau$ by taking derivative twice in the first equation of (9). Then, from (9) and the first equation in (12) we obtain a fourth-order ODE in $k$

$$
\begin{equation*}
k^{(4)}+(5 k-3 \lambda) k^{\prime \prime}+\frac{5}{2}\left(k^{\prime}\right)^{2}+\frac{5}{2} k^{3}-\frac{9}{2} \lambda k^{2}+\left(2 \lambda^{2}+c\right) k+\varepsilon_{1} p_{1}^{2}-c \lambda=0, \tag{16}
\end{equation*}
$$

so that $k^{(4)}$ is a polynomial in $k, k^{\prime}$ and $k^{\prime \prime}$.
Observe that we can identify this Hamiltonian system with solitons (i.e. travelling wave solutions of the form $u(x, t)=u(x-c t))$ of the fifth-order evolution conservative partial differential equation

$$
u_{t}+\left(u_{x x x x}+(5 u-3 \lambda) u_{x x}+\frac{5}{2} u_{x}^{2}+\frac{5}{2} u^{3}-\frac{9}{2} \lambda u^{2}+\left(2 \lambda^{2}+c\right) u\right)_{x}=0,
$$

which is a soliton equation called fifth-order KdV equation (KdV5). The general solution of (16) has been obtained in terms of hyperelliptical functions (see [23]) by separation of the variables of the Hamilton-Jacobi equation in parabolic coordinates.

We do not pretend to find all solutions of this system; however, we will give an explicit solution of a quite interesting particular case when $P_{1}$ and $P_{2}$ are null. In this case, $P_{1}$ and $P_{2}$ should be collinear and so there does not exist a rotational plane. We assume that $P_{2}=\beta P_{1}$, where $\beta$ is a constant. Equalling coefficients of these vector fields we get
(a) $2(k+\lambda) \tau^{\prime}-2 \tau k^{\prime}=\beta\left(k^{\prime \prime}+k^{2}-\lambda k\right)$,
(b) $2 \tau(k-\lambda)=-\beta k^{\prime}$,
(c) $4 \tau^{\prime}=\beta(k-\lambda)$,
(d) $2 k^{\prime \prime}+3 k^{2}-2 \lambda k-\lambda^{2}=2 \beta \tau^{\prime}$.

Taking derivative in (17)(c) and bringing it to the second equation of (9) we deduce that

$$
\begin{equation*}
\tau=-\frac{\beta}{2} \frac{k^{\prime}}{k-\lambda} . \tag{18}
\end{equation*}
$$

Take derivative here and multiply both terms by $k^{\prime} /(k-\lambda)$ to obtain

$$
k^{\prime}=-2 \frac{k^{\prime}}{k-\lambda} \partial_{t}\left(\frac{k^{\prime}}{k-\lambda}\right) .
$$

We may now compute an integral of the latter equation

$$
\begin{equation*}
\left(k^{\prime}\right)^{2}=-(k-\lambda)^{2}(k-d), \tag{19}
\end{equation*}
$$

$d$ being a constant. We take again derivative to get

$$
\begin{equation*}
k^{\prime \prime}=-(k-\lambda)(k-d)-\frac{1}{2}(k-\lambda)^{2} . \tag{20}
\end{equation*}
$$

Finally, combining (18)-(20) it yields (17)(a) and (17)(b). The fourth equation of (17) holds if and only if $d=\frac{\beta^{2}}{4}-\lambda$ holds. Now, from (17)(b) and (19), it is easy to check that

$$
\begin{equation*}
\tau^{2}=-\frac{\beta^{2}}{4}(k-d)=-\frac{\beta^{2}}{4}\left(k+\lambda-\frac{\beta^{2}}{4}\right) . \tag{21}
\end{equation*}
$$

Furthermore, the constant $c$ appearing in (9) is related to $\beta$ and $d$ by

$$
c=\frac{\lambda^{2}}{2}-d^{2}=\frac{\lambda^{2}}{2}-\left(\frac{\beta^{2}}{4}-\lambda\right)^{2} .
$$

Examining Eqs. (19) and (21) we can find explicit solutions depending on both $\lambda$ and $\beta$.
Case $d=\lambda$.
This means that $\beta^{2} / 8=\lambda$ and the solutions are

$$
k(\sigma)=\lambda-\frac{4}{(\sigma+e)^{2}}, \quad \tau(\sigma)^{2}=\frac{8 \lambda}{(\sigma+e)^{2}}=\frac{\beta^{2}}{(\sigma+e)^{2}} .
$$

Case $d>\lambda$.
Now $\beta^{2} / 8>\lambda$ and there are two solutions depending on $k(\sigma) \in(-\infty, \lambda)$ or $k(\sigma) \in(\lambda, d]$

$$
\begin{aligned}
& k(\sigma)=\left(\frac{\beta^{2}}{4}-\lambda\right)-\left(\frac{\beta^{2}}{4}-2 \lambda\right) \operatorname{coth}^{2}\left(\frac{1}{4} \sqrt{\beta^{2}-8 \lambda}(\sigma+e)\right), \\
& \tau(\sigma)^{2}=\frac{\beta^{2}}{4}\left(\frac{\beta^{2}}{4}-2 \lambda\right) \operatorname{coth}^{2}\left(\frac{1}{4} \sqrt{\beta^{2}-8 \lambda}(\sigma+e)\right), \\
& k(\sigma)=\left(\frac{\beta^{2}}{4}-\lambda\right)-\left(\frac{\beta^{2}}{4}-2 \lambda\right) \tanh ^{2}\left(\frac{1}{4} \sqrt{\beta^{2}-8 \lambda}(\sigma+e)\right), \\
& \tau(\sigma)^{2}=\frac{\beta^{2}}{4}\left(\frac{\beta^{2}}{4}-2 \lambda\right) \tanh ^{2}\left(\frac{1}{4} \sqrt{\beta^{2}-8 \lambda}(\sigma+e)\right),
\end{aligned}
$$

respectively.

## Case $d<\lambda$.

Then $\beta^{2} / 8<\lambda$ and the solutions are

$$
\begin{aligned}
& k(\sigma)=\left(\frac{\beta^{2}}{4}-\lambda\right)-\left(2 \lambda-\frac{\beta^{2}}{4}\right) \tan ^{2}\left(\frac{1}{4} \sqrt{8 \lambda-\beta^{2}}(\sigma+e)\right), \\
& \tau(\sigma)^{2}=\frac{\beta^{2}}{4}\left(2 \lambda-\frac{\beta^{2}}{4}\right) \tan ^{2}\left(\frac{1}{4} \sqrt{8 \lambda-\beta^{2}}(\sigma+e)\right) .
\end{aligned}
$$

It is worth investigating how the above particular solutions can be seen as solutions of (16). Set $\varphi=k^{\prime \prime}+\frac{3}{2} k^{2}-$ $3 \lambda k+2 \lambda^{2}+c$ and consider the differential operator

$$
\mathcal{D}=\partial_{\sigma \sigma \sigma}+2 k \partial_{\sigma}+k^{\prime} I
$$

Note that $\mathcal{D}$ is the Jacobi operator of the second Poisson structure of the $K d V$ equation. It is easy to see that

$$
\begin{aligned}
\mathcal{D}(\varphi) & =\varphi^{\prime \prime \prime}+2 k \varphi^{\prime}+k^{\prime} \varphi \\
& =k^{(5)}+(5 k-3 \lambda) k^{\prime \prime \prime}+10 k^{\prime} k^{\prime \prime}+\left(\frac{15}{2} k^{2}-9 \lambda+2 \lambda^{2}+c\right) k^{\prime} \\
& =\partial_{\sigma}\left(k^{(4)}+(5 k-3 \lambda) k^{\prime \prime}+\frac{5}{2}\left(k^{\prime}\right)^{2}+\frac{5}{2} k^{3}-\frac{9}{2} \lambda k^{2}+\left(2 \lambda^{2}+c\right) k+\varepsilon_{1} p_{1}^{2}-c \lambda\right) .
\end{aligned}
$$

It follows that $k$ is a solution of (16) if and only if $\mathcal{D}(\varphi)=0$. Observe that (20) reads

$$
\psi=k^{\prime \prime}+\frac{3}{2} k^{2}-(2 \lambda+d) k+\lambda d+\frac{\lambda^{2}}{2}=0 .
$$

Then we have $\varphi=\psi+\phi$, where $\phi=(d-\lambda)(k-(d+2 \lambda))$, and $\mathcal{D}(\phi)=(d-\lambda) \psi^{\prime}$. We conclude that $\mathcal{D}(\varphi)=\mathcal{D}(\phi)=0$ provided $\psi=0$.

## 4. Solving the natural equations

In order to find the critical null curves of the Lagrangian (7) we have to introduce cylindrical coordinates around the rotational plane $\Pi=\operatorname{span}\left\{P_{1}, P_{2}\right\}$ spanned by $P_{1}$ and $P_{2}$. Such $\Pi=\operatorname{span}\left\{P_{1}, P_{2}\right\}$ can be spacelike, timelike or null depending on the causal character of $P_{1}$ and $P_{2}$. Bearing in mind that $\left\langle P_{1}, P_{2}\right\rangle=0$, in the following table we collect all possibilities.

|  |  |  |  |  |
| :--- | :--- | :--- | :--- | :--- |
|  |  |  |  |  |
| spacelike | timelike | null |  |  |
|  | spacelike | $\Pi$ is spacelike | $\Pi$ is timelike | $\Pi$ is null |
| $P_{2}$ | timelike | $\Pi$ is timelike | never | never |
|  | null | $\Pi$ is null | never | never |

Proposition 7. There exist a suitable translation of the coordinates origin for which the Killing vector field $J$ can be expressed as follows:

$$
\begin{equation*}
J=-P_{1} \wedge P_{2} \wedge \gamma+\varepsilon_{1} \omega P_{1}^{*}=-P_{1} \wedge P_{2} \wedge \gamma-\varepsilon_{1} \varepsilon_{2} p_{2}^{2} P_{1}^{*} \tag{22}
\end{equation*}
$$

where $P_{1}^{*}$ have the same causal character as $P_{1}$ and satisfies that $\left\langle P_{1}, P_{1}^{*}\right\rangle=\varepsilon_{1}$.
Proof. Let $\tilde{\gamma}=\gamma-Y$ be a general translation, $Y$ being a constant vector field. Then we have

$$
J=-P_{1} \wedge P_{2} \wedge \gamma+X=-P_{1} \wedge P_{2} \wedge(\tilde{\gamma}+Y)+X=-P_{1} \wedge P_{2} \wedge \tilde{\gamma}-P_{1} \wedge P_{2} \wedge Y+X
$$

We are going to find vector fields $P_{1}^{*}$ and $Y$ satisfying that $X=\varepsilon_{1} \omega P_{1}^{*}+P_{1} \wedge P_{2} \wedge Y$. To do that we distinguish two cases.
$P_{1}$ is not null. The vector fields $P_{1}$ and $X$ are in $P_{2}^{\perp}$. As $P_{1}$ is not null, we obtain the splitting $P_{2}^{\perp}=\operatorname{span}\left\{P_{1}\right\} \perp P_{1}^{\perp}$, where $P_{1}^{\perp}$ stands for the orthogonal space to $P_{1}$ in $P_{2}^{\perp}$. Therefore, we can find a vector field $Y$ and a constant $\mu$ such that

$$
X=P_{1} \wedge P_{2} \wedge Y+\mu P_{1}
$$

We finish by taking $P_{1}^{*}=\frac{\varepsilon_{1} \mu}{\omega} P_{1}$.
$P_{1}$ is null. Now there exists $P_{1}^{*}$ (not unique) such that $\left\langle P_{1}, P_{1}^{*}\right\rangle=\varepsilon_{1}= \pm 1$ and $P_{2}^{\perp}=P_{1}^{\perp} \oplus \operatorname{span}\left\{P_{1}^{*}\right\}$. As $X \in P_{2}^{\perp}$, then there exists a vector field $Y$ and a constant $\mu$ so that

$$
X=P_{1} \wedge P_{2} \wedge Y+\mu P_{1}^{*}
$$

It follows that $\mu=\varepsilon_{1} \omega$, because $\left\langle X, P_{1}\right\rangle=\left\langle J, P_{1}\right\rangle=\omega$.
Next step is devoted to find suitable expressions in cylindrical coordinates of the critical points of $\mathcal{L}$. The preceding result will be extremely important. We will consider two cases depending on the causal character of the rotational plane $\Pi$.

### 4.1. II is not degenerate

Let $(r, \theta, z, y)$ be the coordinates in $\mathbb{L}^{4}$ given by

$$
\begin{align*}
& X(r, \theta, z, y)=(r \cosh \theta, r \sinh \theta, z, y), \quad r \neq 0  \tag{23}\\
& X(r, \theta, z, y)=(z, r \cos \theta, r \sin \theta, y), \quad r>0, \text { and } \theta \in(0,2 \pi)
\end{align*}
$$

These will be called the non-degenerate cylindrical coordinates around the plane spanned by $\left\{\partial_{y}, \partial_{z}\right\}$. We can assume that $P_{1}$ and $P_{2}$ are collinear with $\partial_{z}$ and $\partial_{y}$, respectively, interchanging $z$ and $y$ if need be. Then, it is no difficult to see that

$$
\left\langle\partial_{\theta}, \partial_{\theta}\right\rangle=-\varepsilon_{1} \varepsilon_{2} r^{2}, \quad\left\langle\partial_{r}, \partial_{r}\right\rangle=1, \quad\left\langle\partial_{z}, \partial_{z}\right\rangle=\varepsilon_{1}, \quad\left\langle\partial_{y}, \partial_{y}\right\rangle=\varepsilon_{2}
$$

vanishing all remaining metric products. Hence, we can write

$$
P_{1}=p_{1} \partial_{z}, \quad P_{2}=p_{2} \partial_{y}, \quad J=-p_{2}\left(p_{1} \partial_{\theta}+\frac{\varepsilon_{1} \varepsilon_{2} p_{2}}{p_{1}} \partial_{z}\right)
$$

which along with $L=r^{\prime} \partial_{r}+\theta^{\prime} \partial_{\theta}+z^{\prime} \partial_{z}+y^{\prime} \partial_{y}$ leads to

$$
\begin{align*}
& \langle J, J\rangle=\varepsilon_{1} j_{2}^{2}\left(\frac{p_{2}^{2}}{p_{1}^{2}}-\varepsilon_{2} p_{1}^{2} r^{2}\right), \quad\left\langle L, P_{1}\right\rangle=\varepsilon_{1} p_{1} z^{\prime}  \tag{24}\\
& \left\langle L, P_{2}\right\rangle=\varepsilon_{2} p_{2} y^{\prime}, \quad\langle L, J\rangle=\varepsilon_{2} p_{2}\left(\varepsilon_{1} p_{1} r^{2} \theta^{\prime}-\frac{p_{2}}{p_{1}} z^{\prime}\right) .
\end{align*}
$$

All these equations yield the following result:

Theorem 8. Let $\gamma: I \longrightarrow \mathbb{L}^{4}$ be a critical point of $\mathcal{L}$ and $\Pi$ a non-null plane. Then $\gamma$ can be described in cylindrical coordinates around $\Pi$ as follows:

$$
\begin{align*}
& r^{2}=\frac{\varepsilon_{2}}{p_{1}^{2}}\left(\frac{p_{2}^{2}}{p_{1}^{2}}-\frac{\varepsilon_{1}\langle J, J\rangle}{p_{2}^{2}}\right), \quad z^{\prime}=\frac{\varepsilon_{1}\left\langle L, P_{1}\right\rangle}{p_{1}}, \\
& y^{\prime}=\frac{\varepsilon_{2}\left\langle L, P_{2}\right\rangle}{p_{2}}, \quad \theta^{\prime}=\frac{1}{p_{1} p_{2} r^{2}}\left(\frac{p_{2}^{2}\left\langle L, P_{1}\right\rangle}{p_{1}^{2}}+\varepsilon_{1} \varepsilon_{2}\langle L, J\rangle\right), \tag{25}
\end{align*}
$$

where

$$
\begin{align*}
& \langle J, J\rangle=64 \tau^{2}\left(\tau^{\prime}\right)^{2}+16\left(\tau k^{\prime}-2(k+\lambda) \tau^{\prime}\right)^{2}+4\left(2 \tau^{2}+\lambda^{2}+2 c\right)\left((k+\lambda)\left(2 \tau^{2}-\lambda^{2}-2 c\right)-8 \lambda \tau^{2}\right), \\
& \left\langle L, P_{1}\right\rangle=-(k-\lambda), \quad\left\langle L, P_{2}\right\rangle=-4 \tau^{\prime}, \quad\langle L, J\rangle=2\left(2 \tau^{2}+\lambda^{2}+2 c\right) . \tag{26}
\end{align*}
$$

### 4.1.1. II is degenerate

We will consider the coordinates $(r, \theta, z, y)$ given by

$$
\begin{equation*}
X(r, \theta, z, y)=\left(z-\frac{\varepsilon r}{2}\left(\theta^{2}+1\right), z-\frac{\varepsilon r}{2}\left(\theta^{2}-1\right),-\varepsilon r \theta, y\right), \quad \text { where } r \in \mathbb{R} \backslash\{0\} \text { and } \theta, z, y \in \mathbb{R}, \tag{27}
\end{equation*}
$$

satisfying

$$
\left\langle\partial_{\theta}, \partial_{\theta}\right\rangle=r^{2}, \quad\left\langle\partial_{y}, \partial_{y}\right\rangle=1, \quad\left\langle\partial_{r}, \partial_{z}\right\rangle=\varepsilon_{1}
$$

vanishing all remaining metric products. As above, we distinguish two subcases depending on the causal character of $P_{1}$.
$P_{1}$ is spacelike. We may assume, without loss of generality, that $P_{1}$ and $\partial_{y}$ are collinear as well as $P_{2}$ and $\partial_{z}$. Then we can write

$$
P_{1}=p_{1} \partial_{y}, \quad P_{2}=b \partial_{z}, \quad J=-b p_{1} \partial_{\theta}
$$

Bearing in mind that $L=r^{\prime} \partial_{r}+\theta^{\prime} \partial_{\theta}+z^{\prime} \partial_{z}+y^{\prime} \partial_{y}$, we deduce

$$
\begin{array}{ll}
\langle J, J\rangle=b^{2} p_{1}^{2} r^{2}, & \left\langle L, P_{1}\right\rangle=p_{1} y^{\prime} \\
\left\langle L, P_{2}\right\rangle=\varepsilon_{1} b r^{\prime}, & \langle L, J\rangle=-b p_{1} r^{2} \theta^{\prime} \tag{28}
\end{array}
$$

$P_{1}$ is null. As above, we may assume that $P_{1}$ and $P_{2}$ are collinear with $\partial_{z}$ and $\partial_{y}$, respectively, and choose $P_{1}^{*}=\left(-\frac{\varepsilon_{1}}{a}, 0,-\frac{\varepsilon_{1}}{a}, 0\right)$. Hence, $P_{1}=a \partial_{z}, P_{2}=p_{2} \partial_{y}$ and we have

$$
\begin{aligned}
& P_{1}^{*}=-\frac{\varepsilon_{1}}{2 a}(\theta-1)^{2} \partial_{z}-\frac{\theta-1}{a r} \partial_{\theta}+\frac{1}{a} \partial_{r}, \\
& J=\frac{p_{2}^{2}}{2 a}(\theta-1)^{2} \partial_{z}+\left(a p_{2}+\frac{\varepsilon_{1} p_{2}^{2}}{a r}(\theta-1)\right) \partial_{\theta}-\frac{\varepsilon_{1} p_{2}^{2}}{a} \partial_{r}
\end{aligned}
$$

An easy computation shows that

$$
\begin{align*}
& \langle J, J\rangle=-p_{2}^{2} r\left(a^{2} r+2 \varepsilon_{1} p_{2}(\theta-1)\right), \\
& 2 a\langle L, J\rangle=-2 p_{2}^{2} z^{\prime}+2 r\left(a^{2} r p_{2}+\varepsilon_{1} p_{2}^{2}(\theta-1)\right) \theta^{\prime}+\varepsilon_{1} p_{2}^{2}(\theta-1)^{2} r^{\prime}, \\
& \left\langle L, P_{1}\right\rangle=\varepsilon_{1} a r^{\prime}  \tag{29}\\
& \left\langle L, P_{2}\right\rangle=p_{2} y^{\prime}, \\
& \langle L, L\rangle=r^{2}\left(\theta^{\prime}\right)^{2}+2 \varepsilon_{1} r^{\prime} z^{\prime}+\left(y^{\prime}\right)^{2} .
\end{align*}
$$

This yields the following result:

Theorem 9. Let $\gamma: I \longrightarrow \mathbb{L}^{4}$ be a critical point of $\mathcal{L}$ and $\Pi$ a degenerate plane. Then $\gamma$ can be described in cylindrical coordinates around $\Pi$ as follows:

| $p_{1}=0, p_{2} \neq 0$ | $p_{1} \neq 0, p_{2}=0$ |
| :--- | :--- |
| $r^{\prime}=\frac{\varepsilon_{1}}{a}\left\langle L, P_{1}\right\rangle$ | $r^{2}=\frac{\langle J, J\rangle}{b^{2} p_{1}^{2}}$ |
| $y^{\prime}=\frac{\left\langle L, P_{2}\right\rangle}{p_{2}}$ | $y^{\prime}=\frac{\left\langle L, P_{1}\right\rangle}{p_{1}}$ |
| $\theta=1-\frac{\varepsilon_{1}}{2 p_{2} r}\left(\frac{\langle J, J\rangle}{p_{2}^{2}}+a^{2} r^{2}\right)$ | $\theta^{\prime}=-b p_{1} \frac{\langle L, J\rangle}{\langle J, J\rangle}$ |
| $z^{\prime}=-\frac{2}{a\left\langle L, P_{1}\right\rangle}\left(r^{2}\left(\theta^{\prime}\right)^{2}+\frac{\left\langle L, P_{2}\right\rangle^{2}}{p_{2}^{2}}\right)$ | $z^{\prime}=-\frac{b}{2\left\langle L, P_{2}\right\rangle}\left(\frac{\left\langle L, P_{1}\right\rangle^{2}}{p_{1}^{2}}+\frac{\langle L, J\rangle^{2}}{\langle J, J\rangle}\right)$ |

where $\langle J, J\rangle,\left\langle L, P_{1}\right\rangle,\left\langle L, P_{2}\right\rangle$ and $\langle L, J\rangle$ are given by Eqs. (28) and (29).
We conclude by declaring that if we were able to integrate the differential equations satisfied by curvatures, then we should find the curve in cylindrical coordinates by quadratures.

## 5. Discussion and outlook

We have studied actions in $(d+1)$-dimensions whose Lagrangians are linear functions on the curvature of the particle path, completing previous works [12,18,19]. We have shown that null helices [20,24] are always possible trajectories of the particles. Otherwise, the non-zero vector field $P_{1}$, obtained from the Euler-Lagrange equation, possesses a nonvanishing space-like component orthogonal to the light-like particle trajectory, which seems to be a manifestation of a generic feature of higher-derivative theories. This vector field can be interpreted as the linear momentum of the particle since it is constant along the curve, which agrees with the conserved linear momentum law. Then, the constants of motion turn out to be the mass and spin of the particle. We have obtained massive, massless and tachyonic states, which correspond to space-like, null and time-like momentum vector $P_{1}$, respectively. Similar results were shown by Plyushchay in [7] for time-like trajectories.

In the simplest geometrical particle model (the constant case) we show that the worldline of the particle is a 3-dimensional Cartan helix whose axis is spanned by the vector $P_{1}$. This was already shown by Nersessian and Ramos using a Hamiltonian formulation, but here we offer an alternative proof which exploits the geometry of the particle path. In the $(d+1)$-dimensional linear case we give the Euler-Lagrange equations (motion equations) of the system. Furthermore, when $d=3$ we describe the nature of the motion equations as a two-dimensional nonrelativistic mechanical system with a cubic potential, showing that is completely integrable in the Liouville sense and solving the Euler-Lagrange equations in a particular case. We also integrate the Cartan equations in cylindrical coordinates around a plane $\Pi$ (which can be interpreted as linear momentum).

To conclude, let us indicate some problems that deserve further attention. First, it would be interesting to introduce techniques to face up to actions in $n=d+1$ dimensions $(d \geq 4)$ whose Lagrangians depend linearly on the curvature and study what kind of trajectories of the relativistic particles appear in this model. Second, one might study actions involving higher order curvatures.

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