# Almost isometries of non-reversible metrics with applications to stationary spacetimes

Miguel Ángel Javaloyes (Universidad de Murcia) (joint work with L. Lichtenfelz and P. Piccione)

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# My collaborators



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## Outline

The talk will consist in three parts:

- 1) Almost isometries of quasi-metrics (abstract setting)
- 2) Almost isometries of Finsler metrics
- 3) Applications to stationary spacetimes (Fermat metrics)

# First part: Almost isometries of quasi-metrics

#### **Definition**

Given a set X, we say that a function  $d: X \times X \to \mathbb{R}$  is a *quasi-metric* if

- (i)  $d(x,y) \ge 0$  for every  $x,y \in X$  and d(x,y) = 0 if and only if x = y,
- (ii)  $d(x,y) + d(y,z) \ge d(x,z)$  (triangle inequality).

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As a consequence of the lack of symmetry, there are two kinds of balls:

- $B_d^+(x,r) = \{ y \in X : d(x,y) < r \}$  (forward balls)
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#### **Definition**

A pair (X, d) will be called a *quasi-metric space* endowed with the topology induced by the family  $B_d^+(x, r) \cap B_d^-(x, r)$ ,  $x \in M$  and r > 0.

Let us observe that this topology coincides with the topology generated by (the balls of) the symmetrized metric  $\widetilde{d}(x,y) = \frac{1}{2}(d(x,y) + d(y,x))$ .

# Quasi-metrics

Quasi-metrics spaces have been studied by many mathematicians:

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Out seminar in the university of Murcia is called "Rey Pastor" after him



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# Quasi-metrics and the triangular function

In a quasi-metric space we can define the length of a continuous curve  $\alpha:[a,b]\subseteq\mathbb{R}\to X$  as

$$\ell(\alpha) = \sup_{\mathcal{P}} \sum_{1=1}^{r} d(\alpha(s_i), \alpha(s_{i+1})),$$

where  $\mathcal{P}$  is the set of partitions  $a = s_1 < s_2 < \ldots < s_{r+1} = b$ ,  $r \in \mathbb{N}$ .

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- We say that  $\alpha$  is *rectifiable* when  $\ell(\alpha)$  is finite.
- Moreover, we say that a curve  $\gamma$  in X from p to q is a *minimizing geodesic* if  $\ell(\gamma) = d(p,q)$ .

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#### Definition

Let us define the *triangular function*  $T: X \times X \times X \to [0, +\infty[$  of a quasi-metric space (X, d) as T(x, y, z) = d(x, y) + d(y, z) - d(x, z) for every  $x, y, z \in X$ .

Evidently, T is continuous.



## Proposition

A curve  $\alpha: [a,b] \subseteq \mathbb{R} \to X$  is a minimizing geodesic of a quasi-metric space (X,d) iff  $T(\alpha(s_1),\alpha(s_2),\alpha(s_3))=0$  for every  $a\leq s_1 < s_2 < s_3 \leq b$ .

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#### **Definition**

Let  $(X_1, d_1)$  and  $(X_2, d_2)$  be two quasi-metric spaces. A bijection  $\varphi: X_1 \to X_2$  is an *almost isometry* if it preserves the triangular function, that is,

$$T_2(\varphi(x), \varphi(y), \varphi(z)) = T_1(x, y, z)$$

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## Corollary

Almost isometries preserve minimizing geodesics.

## Proposition

Given quasi-metric spaces  $(X_1, d_1)$  and  $(X_2, d_2)$ , a bijection  $\varphi : X_1 \to X_2$  is an almost isometry iff  $\exists f : X_2 \to \mathbb{R}$  such that for every  $x, y \in X_1$ :

$$d_2(\varphi(x),\varphi(y)) = d_1(x,y) + f(\varphi(x)) - f(\varphi(y))$$
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#### Proof.

- $\Rightarrow$  (the converse is straightforward)
  - Fix a point  $x_0 \in X_1$  and define  $f: X_2 \to \mathbb{R}$  as  $f(z) = d_2(z, \varphi(x_0)) d_1(\varphi^{-1}(z), x_0)$  for every  $z \in X_2$ .

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  - Given  $x, y \in X_1$ , as  $\varphi$  preserves the triangular function, we have

$$d_1(x,y) + d_1(y,x_0) - d_1(x,x_0) = d_2(\varphi(x),\varphi(y)) + d_2(\varphi(y),\varphi(x_0)) - d_2(\varphi(x),\varphi(x_0)),$$

which is equivalent to (1).

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$$\varphi: (X_1, \widetilde{d}_1) \rightarrow (X_2, \widetilde{d}_2)$$

is an isometry, where

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• Moreover,  $\varphi$  is a homeomorphism and the functions  $f: X_2 \to \mathbb{R}$  are continuous



#### Notation:

- Iso(X, d) is the group of isometries of (X, d)
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## Proposition

- With the above notation,  $\operatorname{Iso}(X, d)$  and  $\operatorname{Iso}(X, d)$  are topological groups endowed with the compact-open topology.
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# Proposition

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- If the topology induced by d is locally compact, then  $\widetilde{\mathrm{Iso}}(X,d)$  and  $\overline{\mathrm{Iso}}(X,d)$  are locally compact.

#### Proof.

The proof follows from the inclusions:

$$\operatorname{Iso}(X,d) \subseteq \widetilde{\operatorname{Iso}}(X,d) \subseteq \operatorname{Iso}(X,\widetilde{d}).$$

#### Local almost isometries

#### Definition

Let  $(X_1, d_1)$  and  $(X_2, d_2)$  be two quasi-metric spaces. A map  $\varphi: X_1 \to X_2$  is a *local almost isometry* if  $\forall x \in X_1$ ,  $\exists \ U \subseteq X_1$ ,  $V \subseteq X_2$  open subsets, with  $x \in U$ , such that  $\varphi|_U: U \to V$  is an almost isometry.

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- define  $d_l$  as the infimum of the lengths of curves between two points. We say that (X, d) is a *length space* when  $d_l = d$ .
- We say that a quasi-metric space is weakly finitely compact if  $B^+(x,r) \cap B^-(x,r)$  are precompact  $\forall x \in X$  and r > 0.

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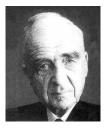
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#### Theorem

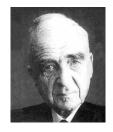
Let  $\varphi:(X_1,d_1)\to (X_2,d_2)$  be a local almost isometry. Assume that  $(X_1,d_1)$  and  $(X_2,d_2)$  are length spaces,  $d_1$  is weakly finitely compact and  $X_2$  is locally arc-connected and simply connected. Then  $\varphi$  is an almost isometry.

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Paul Finsler (1894-1970)

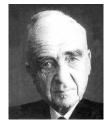
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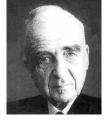
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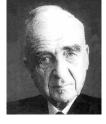


 $g_v(w,z) = \frac{\partial^2}{\partial t \partial s} F(v + tw + sz)^2|_{t=s=0} = \text{Hess}(F^2)_v(w,z)$ 

for every 
$$w, z \in T_{\pi(v)}M$$
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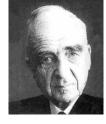
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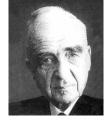
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- Triangle inequality holds in the fibers



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- Named after the norwegian physicist Gunnar Randers (1914-1992):
  - Randers, G.: On an asymmetrical metric in the fourspace of General Relativity. Phys. Rev. (2) **59**, 195–199 (1941)



G. Randers and A. Einstein

### Zermelo metrics

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Geodesics minimize time in the presence of a wind or current W.



### Matsumoto metrics

Given a Riemannian metric g, and a one-form  $\beta$ 

$$M(v) = \frac{g(v, v)}{\sqrt{g(v, v)} - \beta(v)}$$



Sierra Nevada (near Granada)

### Matsumoto metrics

Given a Riemannian metric g, and a one-form  $\beta$ 

$$M(v) = \frac{g(v, v)}{\sqrt{g(v, v)} - \beta(v)}$$

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Geodesics minimize time in the presence of a slope



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Let us define the symmetrized Finsler metric of F as

$$\hat{F}(v) = \frac{1}{2} \big[ F(v) + F(-v) \big]$$

for every  $v \in TM$ . The sum of Finsler metrics is a Finsler metric:



 $\rm M.~A.~J.~AND~M.~S\acute{A}NCHEZ,~\emph{On the definition and examples of}$  Finsler metrics, Arxiv 2011

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#### Lemma

If  $\varphi:(M_1,F_1) \to (M_2,F_2)$  is an almost isometry then

$$\varphi: (M_1, \hat{F}_1) \rightarrow (M_2, \hat{F}_2)$$

is an isometry and  $\varphi$  is smooth.

#### Proof.

- $\bullet$  To see that  $\varphi$  is an isometry prove that preserves the length of curves
- ullet  $\varphi$  is smooth because it is an isometry of a Riemannian average metric

### Proposition

- If  $\exists$  an almost isometry  $\varphi: (M_1, F_1) \to (M_2, F_2)$ , then there exists a smooth  $f: M_2 \to \mathbb{R}$  such that  $\varphi^*(F_1) = F_2 + \mathrm{d}f$ .
- Conversely, if  $\varphi^*(F_1) = F_2 + \mathrm{d}f$ , the map  $\varphi$  is an almost isometry.

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# Proposition

Let (M, F) be a Finsler manifold. Then the extended isometry group  $\widetilde{\mathrm{Iso}}(M, F)$  is a closed subgroup of  $\mathrm{Iso}(M, \hat{F})$ . In particular,  $\widetilde{\mathrm{Iso}}(M, F)$  is a Lie group.

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#### Proof.

Use that  $\widetilde{\mathrm{Iso}}(M,F)\subset \mathrm{Iso}(M,\hat{F})$ 



### Corollary

Let (M,R) be a Randers manifold and  $\varphi:M\to M$  an almost isometry for R. Then  $\varphi$  is an isometry for h.

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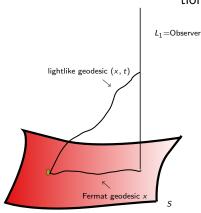
#### Proof.

Just observe that the symmetrized Finsler metric of R is given by  $\hat{R}(v) = \sqrt{h(v,v)}$  for  $v \in TM$ .



# Third part: applications to stationary spacetimes

 $(S \times \mathbb{R}, I)$  is a standard stationary spacetime



S is naturally endowed with a Randers metric F called the Fermat metric

• A spacetime (M, g) is Conformastationary if it admits a timelike Conformal field K, that is, a timelike vector field satisfying

$$\mathcal{L}_{K}g = \lambda g,$$

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ullet Standard Conformastationary means that  $M=S imes\mathbb{R}$  and

$$g((v,\tau),(v,\tau)) = \varphi(g_0(v,v) + 2\omega(v)\tau - \tau^2),$$

in 
$$(x,t) \in S \times \mathbb{R}$$
, where  $(v,\tau) \in T_x S \times \mathbb{R}$ ,  $\varphi : S \times \mathbb{R} \to (0,+\infty)$ 

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A conformastationary spacetime is standard whenever it is distinguishing and the timelike conformal vector field is complete:





M. A. J. AND M. SÁNCHEZ, A note on the existence of standard splittings for conformally stationary spacetimes, Classical Quantum Gravity, 25 (2008), pp. 168001, 7

• First established by Herman Weyl in 1917 for static spacetimes



H. Weyl → (1885-1955)

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- Volker Perlick gave a rigorous proof of this general principle in the same year (1990)



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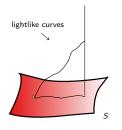


V. Perlick  $\rightarrow$  (Born in 1956)

# Fermat principle in standard stationary spacetimes

 Relativistic Fermat Principle: lightlike pregeodesics are critical points of the arrival time function corresponding to an observer in a suitable class of lightlike curves

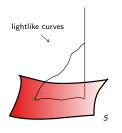
- Relativistic Fermat Principle: lightlike pregeodesics are critical points of the arrival time function corresponding to an observer in a suitable class of lightlike curves
- If you consider as observer  $s \to L_1(s) = (x_1, s)$  in  $(S \times \mathbb{R}, g)$ , given a lightlike curve  $\gamma = (x, t)$ , the arrival time  $\operatorname{AT}(\gamma)$  is





 $t(b) {=} t(a) {+} \textstyle \int_a^b \Bigl( \omega(\dot{x}) {+} \sqrt{g_0(\dot{x}, \dot{x}) {+} \omega(\dot{x})^2} \Bigr) \mathrm{d}s.$ 

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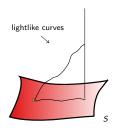




$$t(b)=t(a)+\int_a^b \left(\omega(\dot{x})+\sqrt{g_0(\dot{x},\dot{x})+\omega(\dot{x})^2}\right)\mathrm{d}s.$$

• because  $g_0(\dot{x},\dot{x})+2\omega(\dot{x})\dot{t}-\dot{t}^2=0$   $(g(\dot{\gamma},\dot{\gamma})=0)$ 

- Relativistic Fermat Principle: lightlike pregeodesics are critical points of the arrival time function corresponding to an observer in a suitable class of lightlike curves
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- because  $g_0(\dot{x},\dot{x})+2\omega(\dot{x})\dot{t}-\dot{t}^2=0$   $(g(\dot{\gamma},\dot{\gamma})=0)$
- ullet Let us define the Fermat (Finslerian) metric in S as

$$F(v) = \omega(v) + \sqrt{g_0(v,v) + \omega(v)^2},$$

Pierre de Fermat (1601-1665)

### References



E. Caponio, M. A. J., and A. Masiello, *On the energy functional on Finsler manifolds and applications to stationary spacetimes*, Math. Ann., 351 (2011), pp. 365–392.



E. CAPONIO, M. A. J., AND M. SÁNCHEZ, *On the interplay between Lorentzian Causality and Finsler metrics of Randers type*, Rev. Mat. Iberoamericana, 27 (2011), pp. 919–952.

### References

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#### For a review see:

M. A. J., Conformally standard stationary spacetimes and Fermat metrics, arXiv:1201.1841v1 [math.DG], to appear in Proceedings of GeLoGra 2011.

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#### Theorem

If  $\psi: (S \times \mathbb{R}, g) \to (S \times \mathbb{R}, g)$  is a K-conformal map, then

$$\psi(x,t) = (\varphi(x), t + f(x)),$$

and  $\varphi_*(F) = F + df$  and  $\varphi: (S,h) \to (S,h)$  is an isometry, where

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#### Lemma

 $\operatorname{Conf}_K(M,g)$  (here  $M=\mathbb{R}\times S$ ) is a closed subgroup of  $\operatorname{Conf}(M,g)$ . Moreover the one-parameter subgroup K generated by K is closed and normal in  $\operatorname{Conf}(M,g)$ .

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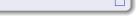
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- Then if  $K^T: M \to M$  is given by  $K^T(x,t) = (x,t+T)$ , it follows that  $\psi \circ K^T \circ \psi^{-1} = K^T$  (K is normal)



### Proposition

The map  $\pi: \mathrm{Conf}_{\mathcal{K}}(M,g) \to \widetilde{\mathrm{Iso}}(S,F)$  defined as  $\pi(\psi) = \varphi$  is a Lie group homomorphism and  $\overline{\pi}: \mathrm{Conf}_{\mathcal{K}}(M,g)/\mathcal{K} \to \widetilde{\mathrm{Iso}}(S,F)$  is an isomorphism.

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- We just have to prove that  $\bar{\pi}$  is one-to-one.
- Injective: if  $\psi_1$  and  $\psi_2$  project on the same almost isometry map  $\varphi$ , then by last Prop.  $\psi_1(x,t)=(\varphi(x),t+f(x)+c_1)$  and  $\psi_2(x,t)=(\varphi(x),t+f(x)+c_2),\ \psi_2\circ\psi_1^{-1}=K^{c_2-c_1}$  and  $[\psi_1]=[\psi_2]$

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- Surjective: given an almost isometry  $\varphi$ , we construct the map

$$\psi(x,t) = (\varphi(x), t + f(x))$$

Clearly, it preserves  $\partial_t$ . By Fermat principle, it maps lightlike pregeodesics to lightlike pregeodesics, then it preserves the lightcone and it must be conformal (by Dajcker-Nomizu [83]).

## **Applications**

### Corollary

Given a manifold S, for a generic set of data  $(g_0, \omega)$ , the stationary metric  $g = g(g_0, \omega)$  on  $M = S \times \mathbb{R}$  has discrete K-conformal group  $\operatorname{Conf}_K(M, g)/\mathcal{K}$ .

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### Corollary

If S is compact, then  $\mathrm{Conf}_K(S \times \mathbb{R}, g)/\mathcal{K}$  and  $\mathrm{Iso}(S, F)$  are compact Lie groups.

## Open problems

- Compute explicitly some extended isometry group
- Which are the Finsler metrics with extended isometry group of maximal dimension?



You can find this talk in http://webs.um.es/imajava

Thanks a lot for this wonderful conference