

Delaunay surfaces of prescribed mean curvature in Nil_3 and $\widetilde{SL}_2(\mathbb{R})$

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Abstract

We obtain a classification result for rotational surfaces in the Heisenberg space and the universal cover of the special linear group, whose mean curvature is given as a prescribed C^1 function depending on their angle function. We show that these surfaces behave like the Delaunay surfaces of constant mean curvature, under some assumptions on the prescribed function. In contrast with the constant mean curvature case, we exhibit the existence of rotational, embedded tori, providing counterexamples of the Alexandrov problem for this class of immersed surfaces.

1 Introduction

The study of surfaces of positive constant mean curvature (CMC surfaces) in the Euclidean space \mathbb{R}^3 is a widely studied topic in the past centuries, whose impact transcends other fields in mathematics such as Analysis of PDE's, Complex Analysis, Geometric Measure Theory and Topology. Among the classical results we highlight the *Delaunay theorem*, that classifies the complete, rotational CMC surfaces as round spheres, cylinders, unduloids and nodoids. In the literature, these surfaces are known as *Delaunay surfaces*.

The theory of CMC surfaces has been extended to further ambient spaces, being of remarkable importance the $\mathbb{E}(\kappa, \tau)$ spaces: the homogeneous, simply connected 3-dimensional manifolds whose isometry group have dimension greater than three and are not space forms. One of the major achievements was the resolution of the Hopf problem by Abresch-Rosenberg [AbRo1, AbRo2], proving that the only immersed CMC spheres are the rotational ones. This milestone attracted the attention of many researches, becoming an active and fruitful field of research; see e.g. [Dan, DHM, FeMi] and references therein for an outline of the development of this theory.

In the past decades, many celebrated results of the theory of CMC surfaces have been generalized to the $\mathbb{E}(\kappa, \tau)$ spaces, among which we remark the classification of complete, rotational CMC surfaces. It was achieved in the product spaces $\mathbb{M}^2(\kappa) \times \mathbb{R}$ by [HsHs, PeRi]; in the Heisenberg space Nil_3 by [Tom]; in the universal cover of the special linear group $\widetilde{SL}_2(\mathbb{R})$ by [Gor, Tor]; and in the Berger spheres \mathbb{S}_b^3 by [Tor]. If the constant mean curvature H_0 satisfies $4H_0^2 + \kappa > 0$, the rotational CMC surfaces behave as the Delaunay surfaces in

Mathematics Subject Classification: 53A10, 53C42, 34C05, 34C40

Keywords: Prescribed mean curvature; Homogeneous 3-spaces; Delaunay-type classification result; Rotational tori.

\mathbb{R}^3 . The value $\sqrt{-\kappa}/2$ for $\kappa \leq 0$ is known as the *critical value* of the mean curvature; there exists a rotational sphere with CMC equal to H_0 if and only if $H_0 > \sqrt{-\kappa}/2$.

Taking as main motivation the Delaunay classification of rotational CMC surfaces in the $\mathbb{E}(\kappa, \tau)$ spaces, in this paper we extend this result to the following class of immersed surfaces:

Definition 1.1 *Let be $\mathfrak{h} \in C^1([-1, 1])$. An immersed, oriented surface Σ in $\mathbb{E}(\kappa, \tau)$ has prescribed mean curvature \mathfrak{h} if its mean curvature function H_Σ is given by*

$$H_\Sigma(p) = \mathfrak{h}(\langle \eta_p, \xi \rangle), \quad \forall p \in \Sigma, \quad (1.1)$$

where η is the unit normal of Σ , ξ is the vertical Killing vector field in $\mathbb{E}(\kappa, \tau)$ and $\langle \eta_p, \xi \rangle =: \nu_p$ is the angle function.

For short, we will say that Σ is an \mathfrak{h} -surface.

Some \mathfrak{h} -surfaces have already appeared in the literature for certain prescribed functions: if $\mathfrak{h} = H_0 \in \mathbb{R}$ we have CMC surfaces and if $\mathfrak{h}(y) = y$ we have translating solitons [Bue1, Bue2, LiMa, Pip]. Also, for a general \mathfrak{h} the author has started the study of global properties of \mathfrak{h} -surfaces in $\mathbb{M}^2(\kappa) \times \mathbb{R}$ [Bue3, Bue4].

The definition of this class of surfaces has its roots in the following *Minkowsky-type prescribed curvature problem* in \mathbb{R}^3 :

Definition 1.2 *Let be $\mathcal{H} \in C^1(\mathbb{S}^2)$. An immersed, oriented surface Σ in \mathbb{R}^3 is an \mathcal{H} -surface if its mean curvature function H_Σ is given by*

$$H_\Sigma(p) = \mathcal{H}(N_p), \quad \forall p \in \Sigma, \quad (1.2)$$

where $N : \Sigma \rightarrow \mathbb{S}^n$ is the Gauss map of Σ .

Again, for $\mathcal{H} = H_0$ a constant we recover the surfaces with constant mean curvature H_0 .

The study of surfaces defined by a prescribed relation between its principal curvatures and its Gauss map goes back, at least, to the famous Minkowski and Christoffel problems for ovaloids [Min, Chr]. When we prescribe the mean curvature as in Equation (1.2), Alexandrov and Pogorelov in the '50s [Ale, Pog] and more recently Guan-Guan [GuGu] and Gálvez-Mira [GaMi1, GaMi2, GaMi3], among others, focused on the existence and uniqueness of immersed \mathcal{H} -spheres. Recently, the author jointly with Gálvez and Mira started to develop the *global theory of hypersurfaces with prescribed mean curvature* in [BGM1, BGM2], taking as starting point the well-studied theory of CMC surfaces in \mathbb{R}^3 . In [BGM1] the authors focused on the study of rotational \mathcal{H} -surfaces, and in [BGM2] global structure results for properly embedded surfaces, including height and curvature estimates, were exhibited.

We highlight the *Delaunay-type classification result* achieved in [BGM1]. Under necessary and sufficient assumptions on the prescribed function, the authors proved that the rotational \mathcal{H} -surfaces behave the same as the CMC Delaunay surfaces in \mathbb{R}^3 . The arbitrariness of the prescribed function made hopeless to find a first integral, even for concrete choices. Instead, the geometric properties of the rotational \mathcal{H} -surfaces were analyzed by means of a *phase plane*

analysis. Inspired by these techniques, the author achieved a Delaunay-type classification result for \mathfrak{h} -surfaces in $\mathbb{M}^2(\kappa) \times \mathbb{R}$ [Bue3].

As happened for \mathcal{H} -surfaces in \mathbb{R}^3 , some hypotheses on \mathfrak{h} are needed if we expect a Delaunay behavior for rotational \mathfrak{h} -surfaces. Throughout this paper, the prescribed function \mathfrak{h} will be always assumed to belong to the following set of functions:

$$\mathfrak{C}^1 := \{\mathfrak{h} \in C^1([-1, 1]); \mathfrak{h}(y) = \mathfrak{h}(-y) > 0 \text{ and } 4\mathfrak{h}(y)^2 + \kappa(1 - y^2) > 0, \forall y \in [-1, 1]\}. \quad (1.3)$$

Note that for the particular case that \mathfrak{h} is a constant $H_0 > 0$ and $\kappa < 0$, the fact that $H_0 \in \mathfrak{C}^1$ reads as $H_0 > \sqrt{-\kappa}/2$, which agrees with the critical value of the mean curvature.

Our main result in this paper is the following Delaunay-type classification result for \mathfrak{h} -surfaces.

Theorem 1.3 *Let be $\mathfrak{h} \in \mathfrak{C}^1$. Up to vertical translations, any complete, rotational \mathfrak{h} -surface in Nil_3 or $\widetilde{SL}_2(\mathbb{R})$ is one of the following:*

1. *The vertical cylinder with radius $x_0 = 2 \left(\sqrt{4\mathfrak{h}(0)^2 + \kappa} + 2\mathfrak{h}(0) \right)^{-1}$ and constant mean curvature $\mathfrak{h}(0)$.*
2. *An embedded \mathfrak{h} -sphere with strictly monotonous angle function.*
3. *A 1-parameter family of properly embedded \mathfrak{h} -unduloids.*
4. *A 1-parameter family of properly immersed \mathfrak{h} -nodoids.*

One of the major issues is that the \mathfrak{h} -nodoids of type 4. in Theorem 1.3 may end up closing. This configuration would lead to the existence of rotational embedded \mathfrak{h} -tori, providing counterexamples to the *Alexandrov problem* for \mathfrak{h} -surfaces. In general, the Alexandrov problem for a class of surfaces asks whether a compact, embedded surface is topologically a sphere. This problem was originally posed by Alexandrov for CMC surfaces in \mathbb{R}^3 and proved by applying his celebrated argument of moving planes.

For CMC surfaces in $\mathbb{E}(\kappa, \tau)$, the Alexandrov problem is fully solved in $\mathbb{M}^2(\kappa) \times \mathbb{R}$ and \mathbb{S}_b^3 . In the Heisenberg space Nil_3 and the space $\widetilde{SL}_2(\mathbb{R})$, the Alexandrov problem is still an outstanding, major open problem. Regarding \mathfrak{h} -surfaces, the Alexandrov problem in $\mathbb{M}^2(\kappa) \times \mathbb{R}$ and \mathbb{S}_b^3 is also solved. The main difference here is that for $\kappa \leq 0$ and $\tau \neq 0$, we prove the existence of rotational embedded \mathfrak{h} -tori for certain prescribed functions $\mathfrak{h} \in \mathfrak{C}^1$.

Theorem 1.4 *There exist rotational and embedded \mathfrak{h} -tori in Nil_3 and $\widetilde{SL}_2(\mathbb{R})$, for certain functions $\mathfrak{h} \in \mathfrak{C}^1$.*

These rotational and embedded \mathfrak{h} -tori provide counterexamples to Alexandrov problem for \mathfrak{h} -surfaces in Nil_3 and $\widetilde{SL}_2(\mathbb{R})$, that is the rotational embedded \mathfrak{h} -spheres given by Item 1. of Theorem 1.3 are not unique in Alexandrov sense.

The arbitrariness of the prescribed function \mathfrak{h} is a major issue when trying to obtain general results regarding \mathfrak{h} -surfaces. In order to adapt our arguments and the original one

exhibited by Tomter [Tom] to the non-existence of rotational \mathfrak{h} -tori, we further assume the following geometric hypothesis: the angle between the E_1 -direction and the tangent vector of the profile curve of the rotational strictly convex \mathfrak{h} -sphere, is a strictly monotonous function. Under this hypothesis, if $\mathfrak{h} \in \mathcal{C}^1$ is decreasing in $[-1, 0]$, we are able to prove the non-existence of rotational \mathfrak{h} -tori.

The rest of the introduction is devoted to further detail the organization of the paper.

In Section 2 we define the $\mathbb{E}(\kappa, \tau)$ spaces as the family of 3-dimensional homogeneous manifolds with a 4-dimensional isometry group, and we introduce a canonical coordinate model. We also define the class of immersed \mathfrak{h} -surfaces in $\mathbb{E}(\kappa, \tau)$ and deduce some properties, making special emphasis on the ambient isometries that are also isometries for \mathfrak{h} -surfaces.

In Section 3 we focus in the analysis of rotational \mathfrak{h} -surfaces. In the same fashion as in [BGM1], the approach will be done by means of a phase plane study of the solutions of the non-linear, autonomous system of ODE's that the coordinates of the profile curve satisfy. In Section 3.1 we deduce the formulas that the profile curve satisfies, and relate the geometry of this curve with the prescribed mean curvature. In Section 3.2 we define the phase plane and exhibit some of its first elements. Sections 3.3 and 3.4 are devoted to study in detail the local and global properties of this phase plane.

Finally, Section 4 is fully devoted to the proof of Theorem 1.3, while in Section 5 we approach the existence of rotational \mathfrak{h} -tori. Specifically, in Section 5.1 we prove Theorem 1.4, and in Section 5.2 we discuss the non-existence of rotational \mathfrak{h} -tori under further geometric hypotheses on \mathfrak{h} .

Acknowledgments: The author is thankful to José A. Gálvez, José M. Manzano and Francisco Torralbo for helpful comments and observations. After revision, the author wants to express his most sincere gratitude to the referee for his/her comments and observations, which led to Section 5 being revised.

2 Immersed \mathfrak{h} -surfaces in the $\mathbb{E}(\kappa, \tau)$ spaces

2.1 The $\mathbb{E}(\kappa, \tau)$ spaces. Consider a homogeneous, simply connected, 3-dimensional manifold whose isometry group has dimension greater than 3 and that is not a space form. Then, its isometry group has dimension 4 and is one of the $\mathbb{E}(\kappa, \tau)$ spaces for some $\kappa, \tau \in \mathbb{R}$ such that $\kappa \neq 4\tau^2$, see [Dan]. A change in the orientation in the space changes τ into $-\tau$, hence we will suppose that $\tau \geq 0$ without losing generality.

The $\mathbb{E}(\kappa, \tau)$ spaces admit a Riemannian submersion $\pi : \mathbb{E}(\kappa, \tau) \rightarrow \mathbb{M}^2(\kappa)$ onto the complete, simply connected surface of constant curvature κ . This fibration has a unitary Killing vector field that will be denoted by ξ , whose integral curves are precisely the fibers of the submersion; recall that the fibers are the sets $\pi^{-1}(q)$, $q \in \mathbb{M}^2(\kappa)$. The group of isometries generated by the Killing vector field ξ are the *vertical translations*.

If $\tau = 0$ we recover the product spaces $\mathbb{M}^2(\kappa) \times \mathbb{R}$ and the submersion π is isomorphic to the projection $\mathbb{M}^2(\kappa) \times \mathbb{R} \rightarrow \mathbb{M}^2(\kappa)$. When $\tau > 0$ we get the Heisenberg space Nil_3 for $\kappa = 0$; the Berger spheres $\mathbb{S}_b^3(\kappa, \tau)$ for $\kappa > 0$; and the universal cover of the special linear group, the space $\widetilde{SL}_2(\mathbb{R})$, for $\kappa < 0$. Note that in the Berger spheres, the projection π is isomorphic to the Hopf fibration.

A key feature is that for every $p \in \mathbb{E}(\kappa, \tau)$ there exists a continuous 1-parameter family of orientation preserving isometries leaving pointwise fixed the fiber $\pi^{-1}(\pi(p))$; these isometries will be called *rotations around the axis* $\pi^{-1}(\pi(p))$.

Next we describe a coordinate model for the $\mathbb{E}(\kappa, \tau)$ spaces; when $\kappa \leq 0$ the model is global, and when $\kappa > 0$ the model is homeomorphic to the universal cover of the space minus one fiber. The notation used is inspired by Section 2.1. in [GaMi3]. We consider $\mathcal{R}(\kappa, \tau)$ to be the space \mathbb{R}^3 if $\kappa \geq 0$, or the disk $\mathbb{D}(2/\sqrt{-\kappa})$ if $\kappa < 0$, endowed with coordinates (x, y, z) , and the metric

$$\langle \cdot, \cdot \rangle = \lambda^2(dx^2 + dy^2) + (\lambda\tau(ydx - xdy) + dz)^2, \quad \lambda = \frac{4}{4 + \kappa(x^2 + y^2)}. \quad (2.1)$$

Then, $\mathcal{R}(\kappa, \tau)$ is isometric to the corresponding $\mathbb{E}(\kappa, \tau)$ space, and the Riemannian submersion is isomorphic to the projection onto the first two coordinates. The vector fields

$$E_1 = \frac{1}{\lambda}\partial_x - \tau y\partial_z, \quad E_2 = \frac{1}{\lambda}\partial_y + \tau x\partial_z, \quad E_3 = \xi = \partial_z,$$

are an orthonormal frame. The E_3 -axis is defined to be the fiber $\pi^{-1}(\pi((0, 0, 0)))$. In this coordinate model we have that the usual rotations

$$(x, y, z) \longmapsto (x \cos \theta + y \sin \theta, -x \sin \theta + y \cos \theta, z), \quad \theta \in \mathbb{R},$$

are the rotations around the E_3 -axis.

Given an immersed surface Σ in $\mathbb{E}(\kappa, \tau)$ and $\eta : \Sigma \rightarrow T\mathbb{E}(\kappa, \tau)$ a unit normal along Σ , the function defined by

$$\nu : \Sigma \rightarrow \mathbb{R}, \quad \nu_p = \langle \eta_p, E_3 \rangle \quad (2.2)$$

is the *angle function* of Σ .

2.2 Isometries and \mathfrak{h} -surfaces. As stated in Definition 1.1 in the Introduction, given $\mathfrak{h} \in C^1([-1, 1])$ an \mathfrak{h} -surface in an $\mathbb{E}(\kappa, \tau)$ space is an immersed surface Σ whose mean curvature satisfies $H_\Sigma(p) = \mathfrak{h}(\nu_p)$, for every $p \in \Sigma$.

Next we describe the isometries in the $\mathbb{E}(\kappa, \tau)$ spaces that send \mathfrak{h} -surfaces into \mathfrak{h} -surfaces. As a matter of fact, translations and rotations around the fibers leave invariant the angle function at any $\mathbb{E}(\kappa, \tau)$ space, hence in virtue of Equation (1.1) send \mathfrak{h} -surfaces into \mathfrak{h} -surfaces.

Note that reflections with respect to horizontal planes if $\tau = 0$, and rotations of angle π around horizontal geodesics if $\tau \neq 0$, change the value ν into $-\nu$. Hence, if \mathfrak{h} is an even function, Σ is an \mathfrak{h} -surface and Ψ is any of those isometries, then $\Psi(\Sigma)$ also satisfies Equation (1.1) and so Ψ send \mathfrak{h} -surfaces into \mathfrak{h} -surfaces.

2.3 Existence of radial solutions. In this section we show the existence of radial solutions for graphical \mathfrak{h} -surfaces defined over a disk with small enough radius. The following proposition is a consequence of a more general existence result for radial solutions of a fully non-linear PDE in the $\mathbb{E}(\kappa, \tau)$ spaces, see Lemma 4.1 in [GaMi3].

Proposition 2.1 *Let be $\mathfrak{h} \in C^1([-1, 1])$. There exists $\delta > 0$ and a function $f : [0, \delta] \rightarrow \mathbb{R}$ such that the radial graph over the distance disk $D(\mathbf{o}, \delta)$ in $\mathbb{M}^2(\kappa)$,*

$$\Sigma_f := \{(x \cos \theta, x \sin \theta, f(x)); x \in [0, \delta], \theta \in [0, 2\pi]\}, \quad f'(0) = 0,$$

with upwards orientation is an \mathfrak{h} -surface in $\mathcal{R}(\kappa, \tau)$. Moreover, Σ_f is unique among graphical \mathfrak{h} -surfaces over $D(\mathbf{o}, \delta)$ having constant boundary data. The same holds for downwards orientation.

3 Rotational \mathfrak{h} -surfaces in homogeneous 3-spaces

3.1 Basic formulas. We begin by locally parametrizing a rotational \mathfrak{h} -surface in the coordinate model $\mathcal{R}(\kappa, \tau)$. Let $\alpha(u) = (x(u), 0, z(u))$ be a curve in the xz -plane². The map

$$\psi(u, \theta) = (x(u) \cos \theta, x(u) \sin \theta, z(u)),$$

defines an immersed surface Σ as the image of $\alpha(u)$ under the rotations of $\mathcal{R}(\kappa, \tau)$ that leave the E_3 -axis pointwise fixed. Note that the projection $\pi(\psi(u, \theta))$ lies in the section $\{z = 0\} \subset \mathcal{R}(\kappa, \tau)$. For $\kappa \geq 0$ this reads as $x(u) \in (0, \infty)$ (since the section $\{z = 0\}$ is \mathbb{R}^2), but for $\kappa < 0$ this implies that $x(u) \in (0, 2/\sqrt{-\kappa})$.

The angle function of Σ in this model is

$$\nu = \frac{4x'}{\sqrt{16(1 + \tau^2 x^2)x'^2 + z'^2(4 + \kappa x^2)^2}}.$$

and the mean curvature H_Σ has the following expression

$$2H_\Sigma = \frac{(4 + \kappa x^2)^2 (z'^3 (16 - \kappa^2 x^4) - 16z'(\tau^2 x^3 x'' + x x'' - x'^2) + 16z'' x x' (1 + \tau^2 x^2))}{4x \left(z'^2 (4 + \kappa x^2)^2 + 16x'^2 (1 + \tau^2 x^2) \right)^{3/2}}. \quad (3.1)$$

Now we consider the metric

$$d\sigma^2 = (1 + \tau^2 x^2)dx^2 + \frac{(4 + \kappa x^2)^2}{16}dz^2 \quad (3.2)$$

in the xz -plane and the arc-length parameter s of α with respect to this metric. A straightforward computation shows that, with this arc-length parameter, the angle function is $\nu = x'$ and the mean curvature is

$$2\varepsilon H_\Sigma = \frac{x(-x''(4 + \kappa x^2)(1 + \tau^2 x^2) + x x'^2(\kappa - 8\tau^2) - \kappa x) - 4x'^2 + 4}{4x\sqrt{1 - x'^2(1 + \tau^2 x^2)}}, \quad \varepsilon := \text{sign}(z').$$

From now on we suppose that Σ is an \mathfrak{h} -surface for some $\mathfrak{h} \in C^1([-1, 1])$, that is $H_\Sigma(p) = \mathfrak{h}(\nu_p)$, $\forall p \in \Sigma$. Solving this equation for x'' yields

$$x'' = \frac{4 - \kappa x^2 - x'^2(4 - x^2(\kappa - 8\tau^2)) - 8\varepsilon x \mathfrak{h}(x')\sqrt{1 - (1 + \tau^2 x^2)x'^2}}{x(4 + \kappa x^2)(1 + \tau^2 x^2)}.$$

²The xz -plane is just the subset $\{y = 0\}$ in $\mathcal{R}(\kappa, \tau)$. Only when $\tau = 0$ it is a totally geodesic surface in $\mathbb{E}(\kappa, \tau)$ isometric to \mathbb{R}^2 .

After the change of variable $x' = y$, this equation transforms into the first order, autonomous system

$$\begin{pmatrix} x \\ y \end{pmatrix}' = \begin{pmatrix} y \\ \frac{4 - \kappa x^2 - y^2(4 - x^2(\kappa - 8\tau^2)) - 8\varepsilon x \mathfrak{h}(y) \sqrt{1 - (1 + \tau^2 x^2) y^2}}{x(4 + \kappa x^2)(1 + \tau^2 x^2)} \end{pmatrix}. \quad (3.3)$$

From the arc-length condition $(1 + \tau^2 x^2)x'^2 + (4 + \kappa x^2)^2/16z'^2 = 1$ we have that the angle function $y = x'$ satisfies

$$\frac{-1}{\sqrt{1 + \tau^2 x^2}} \leq y \leq \frac{1}{\sqrt{1 + \tau^2 x^2}},$$

with equality if and only if the height function z of α has a local extremum. This implies that system (3.3) is only defined for points (x_0, y_0) such that $x_0 > 0$ and $y_0^2 \leq 1/(1 + \tau^2 x_0^2)$. For instance, note that for the case $\tau > 0$, the angle function satisfies $y = \pm 1$ if and only if $x = 0$, which only happens at the axis of rotation.

3.2 The phase plane. Hereinafter we suppose that $\tau > 0$ and $\kappa \leq 0$, i.e. we focus on the spaces Nil_3 and $\widetilde{SL}_2(\mathbb{R})$.

The phase plane of Equation (3.3) is defined as the set

$$\Theta_\varepsilon := \left\{ (x, y); x > 0 \text{ and } y^2 < \frac{1}{1 + \tau^2 x^2} \right\},$$

with coordinates (x, y) denoting the distance to the axis of rotation and the angle function. Note that in virtue of the models introduced in Section 2.1, if $\kappa < 0$ the x -coordinate is defined for $0 < x < 2/\sqrt{-\kappa}$, while if $\kappa = 0$ the x -coordinate is defined for every $x > 0$.

The boundary of Θ_ε consists of the segment $\{0\} \times [-1, 1]$ and the vertical graphs $y = \pm 1/\sqrt{1 + \tau^2 x^2}$. We will denote by Ω^+ (resp. Ω^-) to the component $y = 1/\sqrt{1 + \tau^2 x^2}$ (resp. to the component $y = -1/\sqrt{1 + \tau^2 x^2}$), and by $\Omega := \Omega^+ \cup \Omega^-$.

The orbits are the solutions of system (3.3) and will be denoted by $\gamma(s) = (x(s), y(s))$. The existence and uniqueness of the Cauchy problem associated to Equation (3.3) has as consequence two important facts: *i*) two different orbits cannot intersect in Θ_ε , and *ii*) the orbits are a foliation of Θ_ε by regular C^1 curves.

Although we will remind it in the statement of the main results, **\mathfrak{h} will be always supposed to lie in the space \mathfrak{C}^1** , see Equation (1.3) for a definition of the space \mathfrak{C}^1 .

For example, the fact that \mathfrak{h} is even has the following consequence on the phase plane Θ_ε : if $\gamma(s) = (x(s), y(s))$ is a solution to Equation (3.3), so it is $\tilde{\gamma}(s) = (x(-s), -y(-s))$. This condition is related with the fact that rotations of angle π around horizontal geodesics if $\tau > 0$, are isometries for \mathfrak{h} -surfaces.

Since $\mathfrak{h} \in \mathfrak{C}^1$, a trivial solution to system (3.3) in Θ_1 is the one given by the constant orbit

$$x(s) = \frac{2}{\sqrt{4\mathfrak{h}(0)^2 + \kappa + 2\mathfrak{h}(0)}}, \quad y(s) = 0. \quad (3.4)$$

This point is the *equilibrium* of (3.3) and will be denoted by e_0 . This equilibrium generates a vertical, circular cylinder with constant mean curvature equal to $\mathfrak{h}(0)$. In the next section we will see that no equilibria exist in Θ_{-1} .

From Equation (3.3) we see that the points in Θ_ε with $y'(s) = 0$ are the ones lying in the intersection of Θ_ε with the (possibly disconnected) horizontal graph:

$$x = \Gamma_\varepsilon(y) := 2\sqrt{\frac{1 - y^2}{\kappa(1 - y^2) + 8(\mathfrak{h}(y))^2 + \tau^2 y^2 + 4\varepsilon\mathfrak{h}(y)\sqrt{4\mathfrak{h}(y)^2 + \kappa(1 - y^2) + 4\tau^2 y^2}}}. \quad (3.5)$$

We define $\Gamma_\varepsilon := \{x = \Gamma_\varepsilon(y)\} \cap \Theta_\varepsilon$. Note that the points lying in Γ_ε correspond to points whose angle function has vanishing derivative, and that $\Gamma_1 \cap \{y = 0\} = e_0$ and $\Gamma_{-1} \cap \{y = 0\} = e_{-1}$ for $\kappa > 0$. Again, since \mathfrak{h} is even we get that Γ_ε is symmetric with respect to the axis $y = 0$.

Observation 3.1 *We must clarify the difference between existing an equilibrium point in Θ_{-1} and the point given by $(\Gamma_{-1}(0), 0)$. Even though the latter can exist (for the case that $\Gamma_{-1}(0)$ is well defined), the point $(\Gamma_{-1}(0), 0)$ may not lie in Θ_{-1} . For instance, if $\kappa < 0$ we will see that $\Gamma_{-1}(0) > 2/\sqrt{-\kappa}$, hence it lies outside Θ_{-1} .*

3.3 The structure of Γ_ε . As revealed in the study made in [BGM1], the curve Γ_ε deeply governs the behavior of the orbits in Θ_ε . We study next the properties of the curve Γ_ε in the different phase planes of the spaces Nil_3 and $\widetilde{SL}_2(\mathbb{R})$.

First, recall that the range of the x -coordinate in Θ_ε depends on the value of κ . For the case that $\kappa = 0$, the phase plane Θ_ε is defined for every $x > 0$. For $\kappa < 0$, Θ_ε is defined for $0 < x < 2/\sqrt{-\kappa}$. Because of the restrictions on x for $\kappa < 0$, the curve Γ_ε can leave Θ_ε , or even not exist. Nonetheless, we will see that Γ_1 is always contained in Θ_1 .

Claim 1. The curve Γ_1 for $\kappa < 0$ lies entirely in Θ_1 , i.e. $\Gamma_1(y) < 2/\sqrt{-\kappa}$, $\forall y \in [-1, 1]$.

Proof of the claim : From (3.5) we see that $\Gamma_1(y) < 2/\sqrt{-\kappa}$ if and only if

$$-\kappa(1 - y^2) < \kappa(1 - y^2) + 8(\mathfrak{h}(y))^2 + \tau^2 y^2 + 4\mathfrak{h}(y)\sqrt{4\mathfrak{h}(y)^2 + \kappa(1 - y^2) + 4\tau^2 y^2}.$$

Simplifying we arrive to

$$0 < 4\mathfrak{h}(y)^2 + \kappa(1 - y^2) + 4\tau^2 y^2 + 2\mathfrak{h}(y)\sqrt{4\mathfrak{h}(y)^2 + \kappa(1 - y^2) + 4\tau^2 y^2}.$$

Since $\mathfrak{h} \in \mathcal{C}^1$ in particular $4\mathfrak{h}(y)^2 + \kappa(1 - y^2) > 0$ holds for every $y \in [-1, 1]$, hence the above inequality yields. \square

In any case, the curve Γ_1 is a connected, compact arc in Θ_1 and having the points $(0, \pm 1)$ as endpoints.

Now, we focus on the curve Γ_{-1} in the different spaces. The proof of the following claim holds immediately by just substituting in Equation (3.5) the value $\kappa = 0$.

Claim 2. The curve Γ_{-1} for $\kappa = 0$ is a disconnected bi-graph over the axis $y = 0$, having the points $(0, \pm 1)$ as endpoints and an asymptote at $y = 0$.

As we pointed out, for $\kappa < 0$ the curve Γ_{-1} may leave the phase plane Θ_{-1} . The following claim reveals that the point $(\Gamma_{-1}(0), 0)$ is well defined, and lies outside Θ_{-1} .

Claim 3. The value $\Gamma_{-1}(0)$ for $\kappa = -1$ is well defined and $\Gamma_{-1}(0) > 2/\sqrt{-\kappa}$.

Proof of the claim : Substituting $\Gamma_{-1}(0)$ in Equation (3.5) and simplifying, we get

$$\Gamma_{-1}(0) = \frac{2}{2\mathfrak{h}(0) - \sqrt{4\mathfrak{h}(0)^2 + \kappa}}.$$

Since $\kappa < 0$, the denominator in the above fraction is always well-defined and so it is the value $\Gamma_{-1}(0)$. For proving that $\Gamma_{-1}(0) > 2/\sqrt{-\kappa}$, after a similar computation as in Claim 1. we get

$$0 > 4\mathfrak{h}(0)^2 + \kappa - 2\mathfrak{h}(0)\sqrt{4\mathfrak{h}(0)^2 + \kappa} = \sqrt{4\mathfrak{h}(0)^2 + \kappa} \left(\sqrt{4\mathfrak{h}(0)^2 + \kappa} - 2\mathfrak{h}(0) \right).$$

This time, since $\kappa < 0$ we have that the latter expression is negative, concluding the proof. \square

By the previous claims, in the phase plane Θ_{-1} of both Nil_3 and $\widetilde{SL}_2(\mathbb{R})$, the curve Γ_{-1} is disconnected, has as endpoints $(0, \pm 1)$ and does not intersect the axis $y = 0$. In Nil_3 , Γ_{-1} converges to $y = 0$ and in $\widetilde{SL}_2(\mathbb{R})$, Γ_{-1} leaves Θ_{-1} before intersecting the axis $y = 0$. In particular, no equilibria exist in Θ_{-1} .

In both spaces, the curve Γ_ε and the axis $y = 0$ divide Θ_ε into four connected components, which we will call *monotonicity regions*, where the coordinates $(x(s), y(s))$ of any orbit are monotonous functions; see Figure 1. Hence the behavior of an orbit is uniquely determined by the monotonicity region where it belongs. We describe this behavior next.

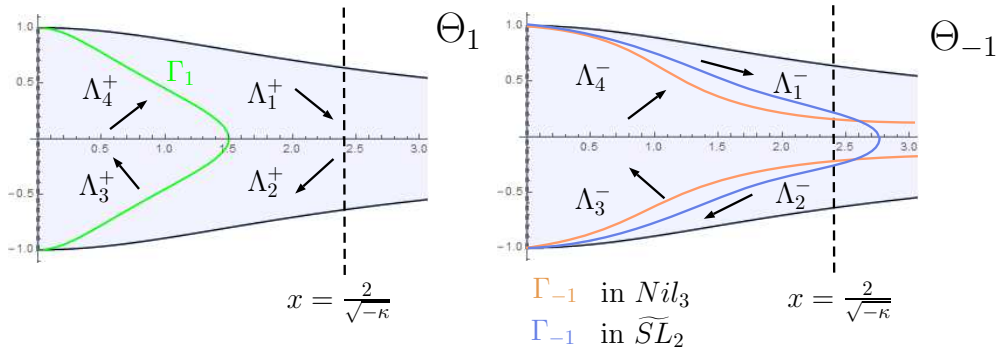


Figure 1: The phase planes Θ_ε , $\varepsilon = \pm 1$ and their monotonicity regions. Note that in $\widetilde{SL}_2(\mathbb{R})$, Θ_ε is only defined for $x < 2/\sqrt{-\kappa}$. The arrows indicate the motion of an orbit.

Proposition 3.2 Let be $(x_0, y_0) \in \Theta_\varepsilon$ and consider an orbit $\gamma(s) = (x(s), y(s))$ such that $\gamma(s_0) = (x_0, y_0)$. Then, the following properties hold:

1. If $y_0 = 0$, then γ is orthogonal to the axis $y = 0$. If $y_0 \neq 0$, then we can see $\gamma(s)$ locally around $\gamma(s_0)$ as a graph $y(x)$. Then:
2. If $x_0 > \Gamma_\varepsilon(y_0)$ (resp. $x_0 < \Gamma_\varepsilon(y_0)$) and $y_0 > 0$, then $y(x)$ is strictly decreasing (resp. increasing) at x_0 .
3. If $x_0 > \Gamma_\varepsilon(y_0)$ (resp. $x_0 < \Gamma_\varepsilon(y_0)$) and $y_0 < 0$, then $y(x)$ is strictly increasing (resp. decreasing) at x_0 .
4. If $x_0 = \Gamma_\varepsilon(y_0)$, then $y'(x_0) = 0$ and $y(x)$ has a local extremum at x_0 .

Proof. First, we study how an orbit intersects the axis $y = 0$. Suppose that $\varepsilon = 1$ and let $\gamma(s) = (x(s), y(s))$ be an orbit in Θ_1 such that $\gamma(0) = (x_0, 0)$ for $x_0 > 0$. Moreover, suppose that $(x_0, 0)$ is not the equilibrium point. From Equation (3.3) we get

$$\gamma'(0) = \left(0, \frac{4 - x_0(8\mathfrak{h}(0) + \kappa x_0)}{x_0(4 + \kappa x_0^2)(1 + \tau^2 x_0^2)} \right).$$

So, $\gamma'(0)$ intersects orthogonally $y = 0$, and it does either upwards or downwards depending on the sign of $p_\kappa(x_0)$, where $p_\kappa(x) = 4 - x(8\mathfrak{h}(0) + \kappa x)$. Note that if $\kappa < 0$, the zeroes of $p_\kappa(x)$ are

$$x_\pm = \frac{2}{2\mathfrak{h}(0) \pm \sqrt{4\mathfrak{h}(0)^2 + \kappa}},$$

and the point $(x_+, 0) \in \Theta_1$ agrees with the equilibrium e_0 defined in (3.4). Depending on the values of κ , this parabola behaves as follows:

- For $\kappa = 0$, $p_\kappa(x)$ reduces to the straight line $y(x) = 4 - 8\mathfrak{h}(0)x$, which has the point $x_+ = 1/(2\mathfrak{h}(0))$ as zero.
- For $\kappa < 0$, $x_-, x_+ > 0$, $x_+ < x_-$, and $p_\kappa(x) > 0$ for $x \in (0, x_+)$. Moreover, $x_+ < 2/\sqrt{-\kappa} < x_-$.

Now, we go back to the study of $\gamma'(0)$. First, suppose that $(x_0, 0)$ lies at the left hand side of e_0 , i.e. $x_0 < x_+$. Hence, $p_\kappa(x_0) > 0$ and thus $\gamma'(0)$ intersects orthogonally the axis $y = 0$ pointing upwards. Analogously, $\gamma'(0)$ points downward whenever $(x_0, 0)$ lies at the right-hand side of e_0 , i.e. when $x_0 > x_+$.

In Θ_{-1} the situation is similar. Again, let $\gamma(s)$ be an orbit in Θ_{-1} such that $\gamma(0) = (x_0, 0)$ for $x_0 > 0$. This time, the value $\gamma'(0)$ is given by

$$\gamma'(0) = \left(0, \frac{4 + x_0(8\mathfrak{h}(0) - \kappa x_0)}{x_0(4 + \kappa x_0^2)(1 + \tau^2 x_0^2)} \right),$$

hence its behavior is determined by the parabola $q_\kappa(x) = 4 + x(8\mathfrak{h}(0) - \kappa x)$. Now, if $\kappa < 0$, $q_\kappa(x)$ has as zeros $z_\pm = -x_\pm$, where x_\pm are the zeroes of $p_\kappa(x)$. Therefore, $q_\kappa(x)$ behaves as follows:

- For $\kappa = 0$, $q_\kappa(x)$ reduces to the straight line $y(x) = 4 + 8\mathfrak{h}(0)x$, which has the point $x = -1/(2\mathfrak{h}(0))$ as zero. In particular, $q_\kappa(x) > 0$, $\forall x > 0$.

- For $\kappa < 0$, $z_-, z_+ < 0$, and $q_\kappa(x) > 0$ for every $x > 0$.

In conclusion, $\gamma'(0)$ points upwards at every $(x_0, 0) \in \Theta_{-1}$.

We have the following consequence in Θ_1 : let $(x_0, 0)$ be a point at the left-hand side of e_0 , and choose $r > 0$ such that $D((x_0, 0), r) \cap \Gamma_1 = \emptyset$. Let $\gamma(s)$ be an orbit in $D((x_0, 0), r)$ such that $\gamma(0) = (x_0, 0)$. Then, the sign of $y'(s)$ is constant and equal to the sign of $y'(0)$, which is positive. By connectedness, the sign of $y'(s)$ of any orbit in $\Lambda_3^+ \cup \Lambda_4^+$ is also positive. The same holds for an orbit contained in $\Lambda_1^+ \cup \Lambda_2^+$, but this time $y'(s)$ is negative. A similar situation holds in Θ_{-1} in the monotonicity regions Λ_3^- and Λ_4^- if $\tau > 0$; we have $y'(s) > 0$ for any orbit contained in those regions.

Finally, we prove that $y'(s) < 0$ for an orbit contained in Λ_1^- or Λ_2^- in Θ_{-1} if $\tau > 0$. By symmetry of the phase plane w.r.t. the axis $y = 0$ it only suffices to prove it in the region Λ_1^- . Fix a point $(x_\infty, y_\infty) \in \Omega^+$ and let $\gamma(s_n) = (x(s_n), y(s_n)) = (x_n, y_n)$ be a sequence of points in an orbit $\gamma(s)$ lying in Λ_1^- and converging to (x_∞, y_∞) . Note that (x_∞, y_∞) are both positive. From Equation (3.3) we see that the values $\gamma'(s_n)$ satisfy:

$$\begin{pmatrix} x'(s_n) \\ y'(s_n) \end{pmatrix} = \begin{pmatrix} y_n \\ \frac{4 - \kappa x_n^2 - y_n^2(4 - x_n^2(\kappa - 8\tau^2)) + 8x_n \mathfrak{h}(y_n) \sqrt{1 - (1 + \tau^2 x_n^2) y_n^2}}{x_n(4 + \kappa x_n^2)(1 + \tau^2 x_n^2)} \end{pmatrix}.$$

Since $(x_\infty, y_\infty) \in \Omega^+$ we get $y_\infty^2(1 + \tau^2 x_\infty^2) = 1$. Taking limits we see that the value $y'(s_n)$ converge to $-(1 - y_\infty^2)(4 + \kappa x_\infty^2)$, which is negative. Because $y'(s) = 0$ only happens at Γ_{-1} , by continuity we get $y'(s) < 0$. Again, a connectedness argument ensures us that this condition is fulfilled for every orbit lying in Λ_1^- , and by symmetry also in Λ_2^- .

This completes the proof of Proposition 3.2. \square

3.4 The structure of the phase plane. In the previous section we focused in how the orbits behave in each monotonicity region, hence how they move through Θ_ε . In this section we exhibit further properties of the phase plane that determine the global and local behavior of an orbit $\gamma(s)$ as it approaches to some point, or tends to *escape* from Θ_ε . These properties are strongly influenced by the underlying geometric problem.

We point out that since e_0 is a solution of Equation (3.3), because $\mathfrak{h} \in C^1$ and by uniqueness of the Cauchy problem, an orbit could converge to e_0 with the parameter $s \rightarrow \infty$. In fact, in [BGM1] we constructed explicit examples converging *directly* to e_0 , that is without spiraling around it. However, this situation cannot happen if \mathfrak{h} is even, as detailed next.

Proposition 3.3 *An orbit γ in Θ_1 cannot converge to e_0 or e_{-1} (for $\kappa > 0$).*

Proof. Let us analyze the structure of the orbits around e_0 . The fact that \mathfrak{h} is even implies that $\mathfrak{h}'(0) = 0$, and the linearized system of Equation (3.3) at e_0 is

$$\begin{pmatrix} u \\ v \end{pmatrix}' = \begin{pmatrix} 0 & 1 \\ F(\kappa, \tau, \mathfrak{h}(0)) & 0 \end{pmatrix} \begin{pmatrix} u \\ v \end{pmatrix},$$

where $F(\kappa, \tau, \mathfrak{h}(0))$ is a negative expression only depending on the mentioned variables; the fact that $\mathfrak{h} \in \mathcal{C}^1$ is key here and in how the element a_{22} in the linearized matrix vanishes.

Hence, the orbits of the linearized system around the origin are ellipses, and by classical theory of non-linear autonomous systems we have two possible configurations around e_0 : either the curves are closed, or they spiral around e_0 , converging to it. However, the latter possibility cannot occur since the phase plane Θ_1 is symmetric with respect to $y = 0$, and e_0 belongs to this axis. In particular, all the orbits in Θ_1 stays at a positive distance from e_0 .

This proof carries over verbatim for the equilibrium point e_{-1} of the phase plane Θ_{-1} when $\kappa > 0$. \square

Next, we analyze the boundary points of Θ_ε that cannot be limit points of orbits.

Proposition 3.4 *An orbit γ cannot converge to a point $(0, y) \in \overline{\Theta_\varepsilon}$, $|y| < 1$.*

Proof. By contradiction, suppose that such an orbit exists and let $\alpha(s) = (x(s), 0, z(s))$ be its associated profile curve. Then, $\gamma(s_n) \rightarrow (0, y)$, $y \in (-1, 1)$, for a sequence s_n . This implies that $x(s_n) \rightarrow 0$ and $x'(s_n) \rightarrow y \in (-1, 1)$, that is $\alpha(s)$ approaches to the E_3 -axis in a non-orthogonal way.

Let Σ be the \mathfrak{h} -surface obtained by rotating α . By the monotonicity properties of the phase plane, we see that a piece of Σ can be written as a graph $z = u(x, y)$ over a punctured disk $D(\mathbf{o}, \delta)$ in $\mathbb{M}^2(\kappa)$. Moreover, the mean curvature $H(x, y)$ viewed as a function in $D(\mathbf{o}, \delta) - \{\mathbf{o}\}$ extends continuously to \mathbf{o} with value $\mathfrak{h}(y)$. By the work of Leandro and Rosenberg [LeRo], we know that Σ extends smoothly to D having vertical unit normal at $u(0, 0)$ and hence angle function equal to ± 1 . This contradicts the fact that $|y| < 1$. \square

In particular, if an \mathfrak{h} -surface intersects the axis of rotation it does in an orthogonal way, i.e. $x' = \pm 1$. Also, Proposition 3.4 ensures us that an orbit $\gamma(s)$ can only converge to points in the boundary of Θ_ε located in Ω . We show that $\gamma(s)$ cannot converge to some $(x_0, y_0) \in \Omega$ for the value of the parameter $s \rightarrow \pm\infty$. Indeed, if $(x(s), y(s)) \rightarrow (x_0, y_0) \in \Omega$ for $s \rightarrow \pm\infty$, then the mean value theorem ensures us that $x'(s) \rightarrow 0$. But $x'(s) = y(s) \rightarrow y_0 \neq 0$, reaching to a contradiction. Hence, if $\gamma(s)$ converges to Ω , reaches it at a finite instant.

Recall that the points $(0, \pm 1)$ are missed by Proposition 3.4. In virtue of Proposition 2.1 we can construct radial \mathfrak{h} -surfaces intersecting orthogonally the axis of rotation over a small enough disk $D(0, \delta)$. These surfaces are defined next.

Definition 3.5 *Let be $\mathfrak{h} \in \mathcal{C}^1$. We define the rotational \mathfrak{h} -surface Σ^+ (resp. Σ^-) as the upwards oriented (resp. downwards oriented) radial solution of Proposition 2.1. Moreover, Σ^+ and Σ^- agree after a rotation around a horizontal geodesic.*

The existence of Σ^+ and Σ^- has the following implication on the phase plane Θ_ε .

Proposition 3.6 *There exists a unique orbit γ_+ (resp. γ_-) in Θ_1 having the point $(0, 1)$ (resp. $(0, -1)$) in $\overline{\Theta_1}$ as endpoint. Moreover, γ_+ and γ_- are symmetric with respect to the axis $y = 0$. There do not exist such orbits in Θ_{-1} .*

Proof. If prove the existence of γ_+ , the existence of γ_- follows from the symmetry of the phase plane Θ_1 .

In virtue of Lemma (2.1), Σ^+ is described as the rotation of a curve $\alpha(s) = (x(s), 0, z(s))$ around the E_3 -axis and such that $x(0) = 0$ and $x'(0) = 1$. Since the mean curvature of Σ^+ at p_0 is positive and Σ^+ is upwards oriented, the mean curvature comparison principle yields $z(s) > 0$ and $z'(s) > 0$ for $s > 0$ small enough. Thus, the orbit $\gamma_+(s)$ generated by $\alpha(s)$ belongs to Θ_1 for $s > 0$ small enough.

The mean curvature comparison principle ensures us that such an orbit cannot exist in Θ_{-1} for upwards oriented graphs. By symmetry of the phase plane we ensure that this condition also holds at the point $(0, -1)$, generating an orbit γ_- that corresponds to the \mathfrak{h} -surface Σ^- . Again, γ_- cannot exist in Θ_{-1} by the mean curvature comparison principle. \square

Thus, the orbit $\gamma_+(s)$ starts at the point $(0, 1) \in \overline{\Theta_1}$, say at the instant $s = 0$, and then is strictly contained in the monotonicity region Λ_1^+ for $s > 0$ small enough. We ensure that $\gamma_+(s)$ cannot intersect again the boundary component $y = 1/\sqrt{1 + \tau^2 x^2}$. In general, we prove that the orbits in Θ_ε cannot have endpoints in Ω arbitrarily.

Proposition 3.7 *Let $\gamma(s)$ be an orbit in Θ_ε and consider $\alpha(s) = (x(s), 0, z(s))$ the associated arc-length parametrized curve. Suppose that $\gamma(s)$ has some $(x_0, y_0) \in \Omega^\pm$ as endpoint at $s = s_0$. Then,*

1. *If $\gamma(s_0) \in \Omega^+$, $z(s)$ has a local minimum at $s = s_0$. In this case, $\gamma(s)$ lies in Θ_1 for $s > s_0$. For $s < s_0$, $\gamma(s)$ belongs to Θ_{-1} .*
2. *If $\gamma(s_0) \in \Omega^-$, $z(s)$ has a local maximum at $s = s_0$. In this case, $\gamma(s)$ lies in Θ_1 for $s < s_0$. For $s > s_0$, $\gamma(s)$ belongs to Θ_{-1} .*

Proof. Suppose that $\gamma_+(s_0) = (x_0, y_0) \in \Omega^+$ for some s_0 and let $\alpha(s) = (x(s), 0, z(s))$ be the arc-length parametrized curve defined by $\gamma(s)$. Note that $x'(s_0) = y_0 > 0$. Because $\gamma(s_0) \in \Omega^+$ we have that the angle function $x'(s) = y(s)$ satisfies $y(s_0) = 1/\sqrt{1 + \tau^2 x(s_0)^2}$, and thus the arc-length condition

$$(1 + \tau^2 x(s)^2)x'(s)^2 + \frac{(4 + \kappa x(s)^2)^2}{16}z'(s)^2 = 1$$

ensures us that $z'(s_0) = 0$. From Equation (3.1) and the fact that $z'(s_0) = 0$, we get

$$2\mathfrak{h}(x'(s_0)) = \frac{(4 + \kappa x(s_0)^2)^2 16x'(s_0) (1 + \tau^2 x(s_0)^2)}{4(16x'(s_0)^2 (1 + \tau^2 x(s_0)^2))^{3/2}} z''(s_0).$$

Since \mathfrak{h} and $x'(s_0) = y(s_0)$ are positive, we see that $z''(s_0)$ is positive as well which yields that $z(s_0)$ is a local minimum of $z(s)$. Thus, $z(s)$ is increasing for $s > s_0$ and decreasing for $s < s_0$. This behavior implies that the orbit describing $\alpha(s)$ lies in Θ_1 for $s > s_0$ and in Θ_{-1}

for $s < s_0$. Hence, this orbit in Θ_{-1} ends at $(x_0, y_0) \in \Omega^+$ and then starts at the same point but this time in Θ_1 .

If $\gamma(s_0) \in \Omega^-$, the proof is similar; just note that $x'(s_0) = y_0 < 0$, hence $z''(s_0)$ is negative. This time, the orbit in Θ_1 ends at Ω^- and then starts at the same point but this time in Θ_{-1} . \square

In any of these situations, i.e. where $z'(s) = 0$ and it changes its monotony, we will say that the orbit in Θ_ε continues in $\Theta_{-\varepsilon}$. and this continuation has to be understood as the extension of the associated \mathfrak{h} -surface having a common point with the same unit normal.

Finally, we focus on whether an orbit can escape from the phase plane Θ_ε .

Proposition 3.8 *Let be $\mathfrak{h} \in \mathfrak{C}^1$ and $\gamma(s) = (x(s), y(s))$ an orbit in Θ_ε . Then $x(s)$ cannot diverge to ∞ if $\kappa = 0$, or tend to $2/\sqrt{-\kappa}$ if $\kappa < 0$.*

Proof. The proof will be done by contradiction and distinguishing the possible values of κ .

Case $\kappa < 0$.

Let $\alpha(s) = (x(s), 0, z(s))$ be the arc-length parametrized curve generated by $\gamma(s)$. Then, the fact that $\alpha(s)$ is the profile of a rotational \mathfrak{h} -surface is equivalent to the following system to be fulfilled

$$\begin{cases} x'(s) = \frac{\cos \theta(s)}{\sqrt{1 + \tau^2 x(s)^2}}, \\ z'(s) = \frac{4 \sin \theta(s)}{4 + \kappa x(s)^2}, \\ \theta'(s) = \frac{1}{(4 + \kappa x(s)^2) \sqrt{1 + \tau^2 x(s)^2}} \left(8\mathfrak{h} \left(\frac{\cos \theta(s)}{\sqrt{1 + \tau^2 x(s)^2}} \right) - \frac{4 - \kappa x(s)^2}{x(s)} \sin \theta(s) \right). \end{cases} \quad (3.6)$$

Here, $\theta(s)$ is the angle that $\alpha'(s)$ makes with the E_1 -direction.

Assume that $\gamma(s)$ lies in Θ_1 with $x(s) \rightarrow 2/\sqrt{-\kappa}$, $s \nearrow s_0$; the case where $s \searrow s_0$ or $\gamma(s)$ lies in Θ_{-1} are proved similarly. Hence, $\gamma(s)$ is contained in Λ_1^+ and stays there as $x(s) \rightarrow 2/\sqrt{-\kappa}$ and $y(s) \rightarrow y_0$, where $y_0 \in [0, 1/\sqrt{1 - 4\tau^2/\kappa}]$. In particular, $\theta(s) \rightarrow \theta_0 \in [0, \pi/2]$, where θ_0 is defined by $\cos \theta_0 = y_0 \sqrt{1 - 4\tau^2/\kappa}$. Because $\mathfrak{h} \in \mathfrak{C}^1$, at the limit $x_0 := 2/\sqrt{-\kappa}$ we have

$$\frac{2}{\sqrt{-\kappa}} \mathfrak{h} \left(\frac{\cos \theta_0}{\sqrt{1 + \tau^2 x_0^2}} \right) > \sqrt{1 - \left(\frac{\cos \theta_0}{\sqrt{1 + \tau^2 x_0^2}} \right)^2} = \sqrt{\frac{\sin^2 \theta_0 + \tau^2 x_0^2}{1 + \tau^2 x_0^2}}.$$

Now, it is trivial to check that the following inequality holds

$$\mathfrak{h} \left(\frac{\cos \theta_0}{\sqrt{1 + \tau^2 x_0^2}} \right) - \frac{\sqrt{-\kappa}}{2} \sin \theta_0 \geq \mathfrak{h} \left(\frac{\cos \theta_0}{\sqrt{1 + \tau^2 x_0^2}} \right) - \frac{\sqrt{-\kappa}}{2} \sqrt{\frac{\sin^2 \theta_0 + \tau^2 x_0^2}{1 + \tau^2 x_0^2}} := c_0 > 0.$$

Taking limit we conclude

$$\lim_{s \rightarrow s_0} 8\mathfrak{h} \left(\frac{\cos \theta}{\sqrt{1 + \tau^2 x^2}} \right) - \frac{4 - \kappa x^2}{x} \sin \theta = 8 \left(\mathfrak{h} \left(\frac{\cos \theta_0}{\sqrt{1 + \tau^2 x_0^2}} \right) - \frac{\sqrt{-\kappa}}{2} \sin \theta_0 \right) \geq 8c_0 > 0.$$

So, the third equation in (3.6) yields that for x close to $2/\sqrt{-\kappa}$ we have $\theta'(s) \geq c_1/(4+\kappa x(s)^2)$, for another positive constant c_1 . In particular, $\lim_{s \rightarrow s_0} \theta'(s) = \infty$.

Now, we see the angle θ as a function $\theta(x)$ of x ; this can be done since $x'(s) > 0$ and by the inverse function theorem. Hence,

$$\theta'(s) = \frac{d\theta}{ds} = \frac{d\theta}{dx} \frac{dx}{ds} = \theta'(x)x'(s) = \theta'(x) \frac{\cos \theta(x)}{\sqrt{1 + \tau^2 x^2}}.$$

In this setting, the third equation in (3.6) for x close enough to $2/\sqrt{-\kappa}$ reads as:

$$\theta'(x) \cos \theta(x) = \frac{1}{4 + \kappa x^2} \left(8\mathfrak{h} \left(\frac{\cos \theta(x)}{\sqrt{1 + \tau^2 x^2}} \right) - \frac{4 - \kappa x^2}{x} \sin \theta(x) \right) \geq \frac{c_1}{4 + \kappa x^2}.$$

Integrating from some fixed, large enough x_0 and x yields

$$\sin \theta(x) \geq \frac{c_1}{2\sqrt{-\kappa}} \operatorname{arctanh} \left(\frac{\sqrt{-\kappa}}{2} x \right) + \sin \theta(x_0).$$

This is a contradiction since the right-hand side tends to infinity as x approaches to $2/\sqrt{-\kappa}$.

Following the idea developed for the case $\kappa < 0$, the case $\kappa = 0$ is proved similarly. Now, the x -coordinate of an orbit $\gamma(s) = (x(s), y(s)) \in \Theta_\varepsilon$ would satisfy $x(s) \rightarrow \infty$. We give a sketch of the proof.

Case $\kappa = 0$. Suppose that $\gamma(s)$ is in the phase plane Θ_1 ; the case Θ_{-1} is analogous. Moreover, we assume that $\gamma(s)$ is strictly contained in Λ_1^+ , since the case when $\gamma(s)$ lies in Λ_2^+ is equivalent by symmetry of Θ_1 . After expressing θ as a function of x , the third equation in (3.6) for x large enough yields

$$\theta'(x) \cos \theta(x) = 2\mathfrak{h} \left(\frac{\cos \theta(x)}{\sqrt{1 + \tau^2 x^2}} \right) - \frac{\sin \theta(x)}{x} \geq c_0 > 0.$$

Explicit integration from some x_0 large enough and $x > x_0$ yields $\sin \theta(x) \geq c_0 x + \sin \theta(x_0)$, which is obviously a contradiction when $x \rightarrow \infty$. \square

4 Proof of Theorem 1.3

We will use the coordinate model $\mathcal{R}(\kappa, \tau)$ as introduced in Section 2. We suppose that the axis of rotation is the E_3 -axis $(0, 0, z)$, $z \in \mathbb{R}$.

The proof will be done by taking advantage from the phase plane analysis made in Section 3. The structure of the phase planes Θ_ε , $\varepsilon = \pm 1$, for the different cases κ, τ was shown in Figure 1 and detailed in Proposition 3.2. In particular, the equilibrium point e_0 in Θ_1 given by Equation (3.4) generates a CMC vertical cylinder, obtaining the first example of the classification.

For the existence of the \mathfrak{h} -sphere, consider the orbit γ_+ starting at the point $(0, 1)$ at the instant $s = 0$ given by Proposition 3.6. For $s > 0$ small enough $\gamma_+(s)$ lies in Λ_1^+ , and from Propositions 3.7 and 3.8 we conclude that $\gamma_+(s)$ is contained in Λ_1^+ until it intersects the axis $y = 0$ at some $\gamma(s_0) = (r_0, 0)$, $r_0 > 0$. This also holds for the orbit $\gamma_-(s)$, i.e. this time

$\gamma_-(s)$ ends at the point $(0, -1)$ at some instant $s_1 > 0$, is contained in Λ_2^+ and comes from intersecting the axis $y = 0$ at a finite instant. Since γ_+ and γ_- are symmetric, they meet orthogonally at the axis $y = 0$. By uniqueness γ_+ and γ_- can be smoothly glued to generate a compact orbit $\gamma_0 := \gamma_+ \cup \gamma_-$ that joins the points $(0, 1)$ and $(0, -1)$.

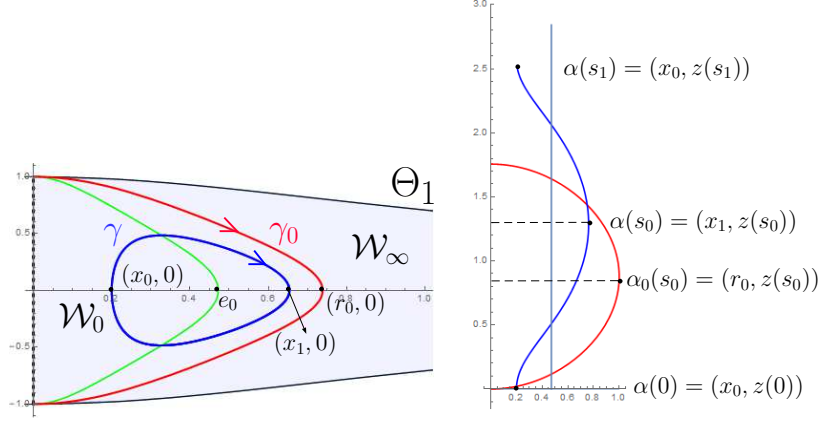


Figure 2: Left: the phase plane Θ_1 and the orbits corresponding to the \mathfrak{h} -sphere, the CMC cylinder and an \mathfrak{h} -unduloid. Right, the profile curves of these \mathfrak{h} -surfaces.

The arc-length parametrized curve $\alpha_0(s)$ associated to $\gamma_0(s)$ intersects the E_3 -axis at the instant $s = 0$, has strictly increasing height function (hence is embedded in $\mathcal{R}(\kappa, \tau)$), and its angle function at the instant $s = s_0$ vanishes; here, the $x(s)$ -coordinate reaches a global maximum and then decreases. By the even condition on \mathfrak{h} , α_0 is symmetric with respect to the horizontal geodesic at height $z(s_0)$ in the xz -plane. The \mathfrak{h} -surface generated by rotating α_0 is an \mathfrak{h} -sphere, denoted by $S_{\mathfrak{h}}$, with strictly decreasing angle function. See Figure 2.

Recall that the orbit γ_0 of the \mathfrak{h} -sphere divides Θ_1 in two connected components: one bounded containing the equilibrium e_0 , that we will denote by \mathcal{W}_0 , and other unbounded that we will denote by \mathcal{W}_∞ .

To prove the existence of the \mathfrak{h} -unduloids, consider $x_0 > 0$ such that $(x_0, 0) \in \mathcal{W}_0$, and suppose that $(x_0, 0)$ is at the left-hand side of e_0 . Let $\gamma(s) = (x(s), y(s))$ be the orbit passing through $(x_0, 0)$ at the instant $s = 0$. Then, for $s > 0$ small enough $\gamma(s)$ is contained in Λ_4^+ and follows its monotonicity direction. By continuity, $\gamma(s)$ has to intersect Γ_1 at some finite point, where the $y(s)$ -coordinate of $\gamma(s)$ reaches a maximum, and then $\gamma(s)$ enters to the region Λ_1^+ . Since $\gamma(s)$ cannot intersect γ_0 by uniqueness of the Cauchy problem and $\gamma(s)$ cannot converge to e_0 with $s \rightarrow \infty$ in virtue of Proposition 3.3, the only possibility for $\gamma(s)$ is to intersect the axis $y = 0$ at some finite point $\gamma(s_0) = (x_1, 0)$ lying at the right-hand side of e_0 . By symmetry of Θ_1 , the same behavior holds in the regions Λ_2^+ and Λ_3^+ and then $\gamma(s)$ reaches again the point $(x_0, 0)$ at some instant $s_1 > 0$, which implies that $\gamma(s)$ is a periodic orbit. Note that in any case, $\gamma(s)$ cannot converge to the segment $\{0\} \times [-1, 1]$ in virtue of Proposition 3.4.

This orbit generates an arc-length parametrized curve $\alpha(s)$ whose height function is strictly increasing (since $\gamma(s) \subset \Theta_1$) and hence $\alpha(s)$ is embedded. The $x(s)$ -coordinate

of $\alpha(s)$ is periodic, its maximum is $x(s_0) = x_1$ and its minimum is $x(0) = x_0$; see Figure 2. The rotation of $\alpha(s)$ generates a properly embedded \mathfrak{h} -surface $U_{\mathfrak{h}}$ that is diffeomorphic to $\mathbb{S}^1 \times \mathbb{R}$ and with periodic distance to the E_3 -axis, that is, $U_{\mathfrak{h}}$ is an \mathfrak{h} -unduloid.

Moreover, each \mathfrak{h} -unduloid is uniquely determined by the value x_0 that agrees with the radius of the smallest circumference contained in $U_{\mathfrak{h}}$. Thus, the family of \mathfrak{h} -unduloids is a continuous family $\{U_{\mathfrak{h}}(r)\}$ parametrized by the *necksize* of their *waists*, where $0 < r < 2(\sqrt{4\mathfrak{h}(0)^2 + \kappa} + 2\mathfrak{h}(0))^{-1}$. Similarly to the CMC case, when $r \rightarrow 2(\sqrt{4\mathfrak{h}(0)^2 + \kappa} + 2\mathfrak{h}(0))^{-1}$ the \mathfrak{h} -unduloids $U_{\mathfrak{h}}(r)$ converge to the vertical cylinder of CMC $\mathfrak{h}(0)$, and when $r \rightarrow 0$ the \mathfrak{h} -unduloids converge to a singular chain of tangent \mathfrak{h} -spheres.

Lastly, we prove the existence of the \mathfrak{h} -nodoids. For that, let $(r_0, 0)$ be the point of intersection of γ_0 with $y = 0$, and fix some $x_0 > r_0$. Consider the orbit $\gamma(s)$ passing through $(x_0, 0)$ at the instant $s = 0$. For $s > 0$ small enough, $\gamma(s)$ lies in Λ_2^+ , and since γ_0 and γ cannot intersect each other, $\gamma(s)$ has some $\gamma(s_0) = (x_1, y_1) \in \Omega^-$, $x_1 > 0$, $y_1 < 0$ as endpoint. By symmetry, $\gamma(s)$ for $s < 0$ has the same behavior at Λ_1^+ , having the point $\gamma(-s_0) = (x_1, -y_1) \in \Omega^+$ as endpoint. This orbit generates an arc-length parametrized curve $\alpha(s)$ which is also symmetric with respect to the rotation of angle π around the horizontal geodesic in the xz -plane at height $z(0)$; after a vertical translation we can suppose that $z(0) = 0$. At this height, the $x(s)$ -coordinate of $\alpha(s)$ reaches its maximum x_0 , and then decreases to the value x_1 . The height $z(s)$ reaches a minimum at $s = -s_0$ where $z'(-s_0) = 0$, then increases and reaches a maximum at $s = s_0$ where again $z'(s_0) = 0$. See Figure 3.

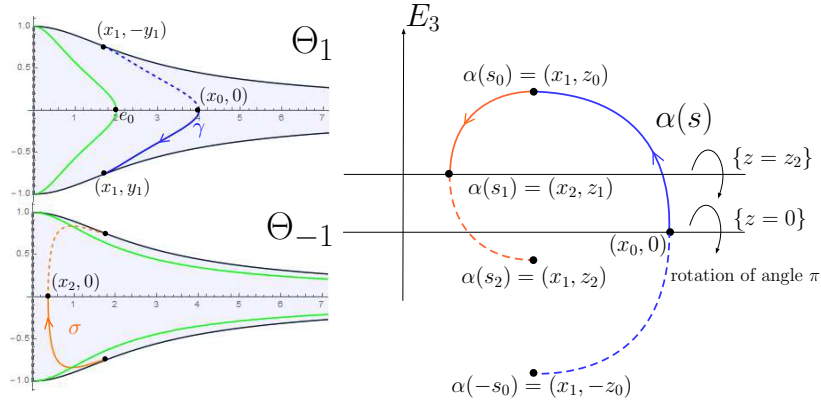


Figure 3: Left: the phase plane and the orbit generating an \mathfrak{h} -nodoid. Right: the profile curve of the \mathfrak{h} -nodoid.

Now, for $s > s_0$ the function $z(s)$ is decreasing, hence $\alpha(s)$ for $s > s_0$ generates an orbit $\sigma(s)$ in Θ_{-1} having the point $(x_1, y_1) \in \Omega^-$ as endpoint. The monotonicity properties of Θ_{-1} and Proposition 3.4 ensures us that $\sigma(s)$ has to intersect the axis $y = 0$ at some finite point $\sigma(s_1) = (x_2, 0)$, $x_2 > 0$. By symmetry, $\sigma(s)$ ends up having the point $(x_1, -y_1) \in \Omega^+$ as endpoint for some $s = s_2$. Thus, the height of $\alpha(s)$ is strictly decreasing starting at the value z_0 and having the value $z(s_2)$ as minimum. This time, the distance $x(s)$ to the E_3 -axis has the value x_2 as minimum. See again Figure 3.

5 Existence of rotational \mathfrak{h} -tori

5.1 Proof of Theorem 1.4. Following the notation introduced in the existence of the \mathfrak{h} -nodoids in the proof of Theorem 1.3, assume that $\alpha(s) = (x(s), 0, z(s))$ is the profile curve of an \mathfrak{h} -nodoid, as represented in Figure 3, whose coordinates and the angle $\theta(s)$ between $\alpha'(s)$ and E_1 satisfy system (3.6). Hence, $\alpha(s)$ is defined for $s \in [0, s_1]$, and $\theta(s)$ satisfies $\theta(0) = \pi/2$, $\theta(s_0) = \pi$ and $\theta(s_1) = 3\pi/2$. Also, for $s \in [0, s_1]$ the orbit that defines $\alpha(s)$ is contained in the region $\{y \leq 0\}$ of the phase plane, hence $y(s) = x'(s) < 0$ for $s \in (0, s_1)$ and $y(0) = y(s_1) = 0$.

First, we prove a technical result concerning the monotonicity of $\theta(s)$ that will be useful in the sequel. Recall that for $s \in [s_0, s_1]$, i.e. when $\theta \in [\pi, 3\pi/2]$, we have $\theta'(s) > 0$ by just substituting in (3.6). When $s \in [0, s_0]$, we prove the following:

Proposition 5.1 *Assume that $\mathfrak{h} \in \mathcal{C}^1$ is increasing in $[-1, 0]$. In the above conditions, $\theta'(s) > 0$ for every $s \in [0, s_0]$.*

Proof. First, we see that $\theta'(0) > 0$ and $\theta'(s_0) > 0$. The latter is immediate by just substituting at (3.6) and bearing in mind that $\theta(s_0) = \pi$. The former holds by differentiating $y = x' = \cos \theta / \sqrt{1 + \tau^2 x^2}$ and substituting at $s = 0$. Indeed,

$$y'(s) = \frac{-\sin \theta(s) \theta'(s) \sqrt{1 + \tau^2 x(s)^2} - \frac{\cos \theta(s) \tau^2 x(s) x'(s)}{\sqrt{1 + \tau^2 x(s)^2}}}{1 + \tau^2 x(s)^2}, \quad (5.1)$$

and at $s = 0$ one has $\theta(0) = \pi/2$, hence the sign of $y'(0)$ and $-\theta'(0)$ agree. Since $y'(0) < 0$ by the properties of the phase plane, we conclude that $\theta'(0) > 0$.

Now, assume that $\theta'(\bar{s}) = 0$ for some $\bar{s} \in (0, s_0)$. We differentiate $\theta'(s)$ in (3.6) and substitute at $s = \bar{s}$, obtaining

$$\theta''(\bar{s}) = \frac{1}{(4 + \kappa x(\bar{s})) \sqrt{1 + \tau^2 x(\bar{s})^2}} (8\mathfrak{h}'(x'(\bar{s}))x''(\bar{s}) - f'(x(\bar{s}))x'(\bar{s}) \sin \theta(\bar{s})),$$

where $f(x) := (4 - \kappa x^2)/x$ is such that $f'(x) < 0$; recall that $f(x)$ is defined for every $x > 0$ if $\kappa = 0$ and for $x \in (0, 2/\sqrt{-\kappa})$ if $\kappa < 0$. Finally, $x''(\bar{s}) = y'(\bar{s}) < 0$ and $x'(\bar{s}) = y(\bar{s}) < 0$ by the monotonicity of the phase plane, $f'(x(\bar{s})) < 0$ and $\mathfrak{h}'(x'(\bar{s})) \geq 0$ by hypothesis, therefore we conclude that $\theta''(\bar{s}) < 0$. As a matter of fact, we have proved that if $\theta'(s)$ vanishes at some point, $\theta(s)$ has a local maximum.

Summarizing, $\theta(s)$ starts at $s = 0$ increasing and reaches $s = s_1$ being also increasing. If $\theta'(\bar{s})$ vanishes at some \bar{s} , it does at a local maximum point and so there should exist another \bar{s}_1 where $\theta(s)$ would attain a local minimum in order to increase again and reach $s = s_1$. This is a contradiction, since we proved that $\theta'(s)$ can only vanish at local maximum points, hence $\theta'(s) > 0$ everywhere in $[0, s_1]$. \square

For a general function $\mathfrak{h} \in \mathcal{C}^1$ we cannot ensure that $\theta'(s) > 0$ for every $s \in [0, s_0]$. Indeed, the particular geometry of Nil_3 and $\widetilde{SL}_2(\mathbb{R})$, induced by the fact that $\tau \neq 0$, has

as implication the existence of functions $\mathfrak{h} \in \mathcal{C}^1$ for which the associated strictly convex³ \mathfrak{h} -sphere has points where θ' vanishes and changes of sign; see Figure 4. In particular, there are \mathfrak{h} -nodoids close enough to the \mathfrak{h} -sphere for which $\theta'(s)$ vanishes.

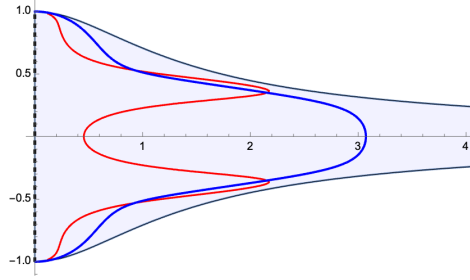


Figure 4: The phase plane for the prescribed function $\mathfrak{h}(y) = 0.1 + \cos(4y)^2$. In red, the curve where $\theta'(s) = 0$. In blue, the orbit of the strictly convex \mathfrak{h} -sphere. Here, the space is Nil_3 .

Now we stand in position to prove the existence of $\mathfrak{h} \in \mathcal{C}^1$ for which there exists a rotational embedded \mathfrak{h} -torus. The fact that $\tau > 0$ has strong implications on the behavior of the angle function of a rotational \mathfrak{h} -surface. This, along with the arbitrariness of the prescribed function \mathfrak{h} , was the key that suggested us that the existence of rotational \mathfrak{h} -tori could be possible in these spaces.

Fix some constant $H_0 \in \mathcal{C}^1$ and $\delta > 0$ small enough (we will determine later a bound for δ). Let be $x_1 = x(s_0)$ and define $\nu_0 := x'(s_0) = -1/\sqrt{1 + \tau^2 x_1^2}$. Given $\lambda > H_0$, consider the following function:

$$\mathfrak{h}_\lambda(y) = \begin{cases} H_0 & \text{if } y \in [-1, \nu_0 - \delta] \\ \lambda & \text{if } y \in [\nu_0, -\nu_0] \\ H_0 & \text{if } y \in [-\nu_0 + \delta, 1], \end{cases}$$

and extend it to be C^1 in the whole interval $[-1, 1]$ in such a way that is increasing in $[-1, 0]$; when $\lambda = H_0$ we choose $\mathfrak{h}_\lambda \equiv H_0$. In particular, \mathfrak{h}_λ lies in the hypothesis of Proposition 5.1.

From system (3.6), the function $z(s)$ can be written by the integral formula $z(s) = \int_0^s 4 \sin \theta(t)/(4 + \kappa x(t)^2) dt$. Because $\theta'(s) > 0$, we can express the functions $x(s)$, $z(s)$ and $\nu(s)$ as functions of θ . The chain rule yields

$$\frac{dz}{d\theta} = \frac{z'(s)}{\theta'(s)} = \frac{4 \sin \theta \sqrt{1 + \tau^2 x^2}}{8\mathfrak{h}(\nu) - \frac{4 - \kappa x^2}{x} \sin \theta}. \quad (5.2)$$

Thus, we have

$$I_1(\mathfrak{h}_\lambda) := - \int_\pi^{3\pi/2} \frac{dz}{d\theta} d\theta = z(s_0) - z(s_1); \quad I_2(\mathfrak{h}_\lambda) := \int_{\pi/2}^\pi \frac{dz}{d\theta} d\theta = z(s_0) - z(0). \quad (5.3)$$

Note that for $\lambda = H_0$ then $I_1(H_0) < I_2(H_0)$, hence $z(s_1) > z(0)$ and no CMC nodoid closes as a rotational CMC torus. See e.g. [Tom], pp. 490, for the proof in Nil_3 . In $\widetilde{SL}_2(\mathbb{R})$ the proof carries over verbatim.

³Recall that by strictly convex we mean that the angle function $x' = \langle \eta, \xi \rangle$ is strictly monotonous.

If we prove that $I_1(\mathfrak{h}_{\lambda_*}) > I_2(\mathfrak{h}_{\lambda_*})$ for some $\lambda_* > H_0$, by continuity we would have $I_2(\mathfrak{h}_{\lambda_0}) = I_1(\mathfrak{h}_{\lambda_0})$ for some $\lambda_0 \in (H_0, \lambda_*)$, and conclude the existence of an \mathfrak{h}_{λ_0} -torus. We prove this next.

The motion of the orbits in the phase plane implies that $\nu(\theta)$ behaves as follows: $\nu(\theta)$ is strictly decreasing for $\theta \in (0, \pi/2)$, and when reaches the value $\theta = \pi$ it keeps decreasing until reaching a global minimum at some $\hat{\theta} \in (\pi, 3\pi/2)$. Then, $\nu(\theta)$ is strictly increasing until reaching the value $\theta = 3\pi/2$, where it vanishes again.

We will denote with a sub-index $(\cdot)_\lambda$ to the functions corresponding to the prescribed function \mathfrak{h}_λ . Let $\gamma_\lambda = (x_\lambda, y_\lambda)$ (resp. γ_{H_0}) be the orbit in Θ_1 for the prescribed function \mathfrak{h}_λ (resp. for H_0) passing through $(x_0, 0)$ at $s = 0$, and assume moreover that γ_{H_0} generates a CMC nodoid. Bearing in mind that $\mathfrak{h}_\lambda \geq H_0$ and $\mathfrak{h}_\lambda(0) > H_0$, a standard comparison argument between the coordinates of γ_λ and γ_{H_0} yields that γ_{H_0} lies at the left-hand side of γ_λ in Θ_1 . Hence, if $\theta \in (\pi/2, \pi)$ then $x_\lambda(\theta) \in (x_1, x_0)$, $\nu_\lambda(\theta) \in (\nu_0, 0)$ and $\mathfrak{h}_\lambda(\nu_\lambda(\theta)) = \lambda$. Since $f(x) = (4 - \kappa x^2)/x$ is positive and strictly decreasing, the following holds

$$I_2(\mathfrak{h}_\lambda) = \int_{\pi/2}^{\pi} \frac{4 \sin \theta \sqrt{1 + \tau^2 x_\lambda^2}}{8\mathfrak{h}_\lambda(\nu_\lambda) - f(x_\lambda) \sin \theta} d\theta < \int_{\pi/2}^{\pi} \frac{4 \sin \theta \sqrt{1 + \tau^2 x_0^2}}{8\lambda - f(x_1) \sin \theta} d\theta.$$

Therefore, $I_2(\mathfrak{h}_\lambda) \rightarrow 0$ as $\lambda \rightarrow \infty$.

On the other hand, let be $\theta_1 < \theta_2 \in (\pi, 3\pi/2)$ such that $\nu_{H_0}(\theta_i) = \nu_0 - \delta$; we choose δ small enough so that the minimum of ν_{H_0} is smaller than $\nu_0 - \delta$, and consequently θ_i are well defined. Hence, we can split $I_1(\mathfrak{h}_\lambda)$ as

$$- \int_{\pi}^{\theta_1} \frac{4 \sin \theta \sqrt{1 + \tau^2 x_\lambda^2}}{8\mathfrak{h}_\lambda(\nu_\lambda) - f(x_\lambda) \sin \theta} d\theta - \int_{\theta_1}^{\theta_2} \frac{4 \sin \theta \sqrt{1 + \tau^2 x_\lambda^2}}{8\mathfrak{h}_\lambda(\nu_\lambda) - f(x_\lambda) \sin \theta} d\theta - \int_{\theta_2}^{3\pi/2} \frac{4 \sin \theta \sqrt{1 + \tau^2 x_\lambda^2}}{8\mathfrak{h}_\lambda(\nu_\lambda) - f(x_\lambda) \sin \theta} d\theta.$$

Note that for $\theta \in (\theta_1, \theta_2)$ in the second integral, the function $\mathfrak{h}_\lambda(\nu_\lambda(\theta))$ is equal to H_0 . In particular, all the geometric quantities appearing are the ones corresponding for the constant function H_0 . Therefore,

$$I_1(\mathfrak{h}_\lambda) > - \int_{\theta_1}^{\theta_2} \frac{4 \sin \theta \sqrt{1 + \tau^2 x_\lambda^2}}{8\mathfrak{h}_\lambda(\nu_\lambda) - f(x_\lambda) \sin \theta} d\theta = - \int_{\theta_1}^{\theta_2} \frac{4 \sin \theta \sqrt{1 + \tau^2 x_{H_0}^2}}{8H_0 - f(x_{H_0}) \sin \theta} d\theta = c_0 > 0,$$

where c_0 is a positive constant that does not depend on λ . Thus, for λ big enough we have $I_1(\mathfrak{h}_\lambda) > I_2(\mathfrak{h}_\lambda)$, hence there is some $\lambda_0 > H_0$ such that $I_1(\mathfrak{h}_{\lambda_0}) = I_2(\mathfrak{h}_{\lambda_0})$, i.e. there exist a rotational embedded \mathfrak{h}_{λ_0} -torus. See Figure 5.

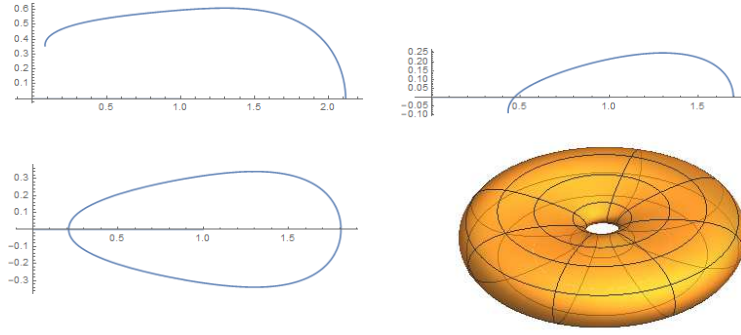


Figure 5: Top: the profile curves for different functions \mathfrak{h}_λ , for which $I_1(\mathfrak{h}_\lambda) - I_2(\mathfrak{h}_\lambda)$ changes sign. Bottom: the profile curve and the rotational \mathfrak{h}_{λ_0} -torus for a function \mathfrak{h}_{λ_0} such that $I_1(\mathfrak{h}_{\lambda_0}) = I_2(\mathfrak{h}_{\lambda_0})$. Here, the space is Nil_3 .

5.2 Non-existence of rotational \mathfrak{h} -tori. We finally discuss the non-existence of rotational \mathfrak{h} -tori for certain functions $\mathfrak{h} \in \mathcal{C}^1$, besides the trivial cases $\mathfrak{h} = H_0 \in \mathbb{R}$. The main difficulty for applying our arguments is that, given some $\mathfrak{h} \in \mathcal{C}^1$, the function $\theta(s)$ of an \mathfrak{h} -nodoid may not be strictly monotonous, hence we cannot express the height $z(s)$ as a function of θ by the relation (5.2) and compare the integrals (5.3). In fact, as we already pointed out in Section 5.1 and showed in Figure 4, there are functions $\mathfrak{h} \in \mathcal{C}^1$ for which the condition $\theta'(s) > 0$ does not hold on the \mathfrak{h} -sphere, hence on \mathfrak{h} -nodoids close enough to the \mathfrak{h} -sphere.

Let us focus on the space Nil_3 , since the computations in this space are easier. From Equation (5.1), we see that

$$-\theta'(s) \sin \theta(s) \sqrt{1 + \tau^2 x^2} = y'(s)(1 + \tau^2 x(s)^2) + \tau^2 x(s)y(s)^2,$$

hence the points in Θ_1 with $\theta'(s) = 0$ are the ones lying in the intersection of Θ_1 with the curve

$$\Lambda := \{(x, y) \in \Theta_1; \Lambda(x, y) := G(x, y)(1 + \tau^2 x^2) + \tau^2 xy^2 = 0\},$$

where $G(x, y)$ is such that $y'(s) = G(x(s), y(s))$ in virtue of (3.3). This curve can be represented as the horizontal graph

$$\Lambda = \{(x, y) \in \Theta_1; x = \Phi(y) := \sqrt{\frac{1 - y^2}{4\mathfrak{h}(y)^2 + \tau^2 y^2}}\}.$$

In the space $\widetilde{SL}_2(\mathbb{R})$, the curve Λ can be also expressed as a horizontal graph, but its expression is way more difficult than the one in Nil_3 . For the sake of clarity, details are omitted. From this expression, we deduce that Λ is a compact arc joining the points $(0, \pm 1)$. Moreover, given an orbit $(x(s), y(s)) \subset \Theta_1$, $x(s) > \Phi(y(s))$ if and only if $\theta'(s) > 0$.

Let us denote by $\gamma_0(s)$ to the orbit of the \mathfrak{h} -sphere, which starts at $\gamma_0(0) = (0, 1)$, intersects the axis $y = 0$ at $\gamma_0(s_0) = (r_0, 0)$, for $r_0 > 0$, and ends at $\gamma_0(s_1) = (0, -1)$. Let $\theta_{S_{\mathfrak{h}}}(s)$ be the angle of between the derivative of its profile curve and the E_1 -direction. Therefore, γ_0 lies

at the right-hand side of Λ at $(r_0, 0)$ since $\theta'_{S_{\mathfrak{h}}}(s_0) > 0$ by substituting in Equation (5.1). If $\theta'_{S_{\mathfrak{h}}}(s) > 0$ for every $s \in (0, s_1)$, then γ_0 lies at the right-hand side of Λ until they both meet at the points $(0, \pm 1)$. Since the orbits of the \mathfrak{h} -nodoids cannot intersect γ_0 by uniqueness, none of these orbits can intersect the curve Λ , hence $\theta'(s) > 0$ for every \mathfrak{h} -nodoid. Bearing these discussions in mind, the following property guarantees that the function $\theta(s)$ of each \mathfrak{h} -nodoid satisfies $\theta'(s) > 0$.

*A function $\mathfrak{h} \in \mathfrak{C}^1$ is said to satisfy the property **(P)** if the function $\theta_{S_{\mathfrak{h}}}(s)$ of the rotational strictly convex \mathfrak{h} -sphere is strictly monotonous.*

We denote by $\mathfrak{C}_{\mathbf{P}}^1$ to the functions of \mathfrak{C}^1 that also satisfy the property **(P)**. The arbitrariness of \mathfrak{h} and the lack of having a first integral, is a major difficulty when trying to explicitly check if property **(P)** holds for such an \mathfrak{h} . Examples of functions $\mathfrak{h} \in \mathfrak{C}_{\mathbf{P}}^1$ are the ones increasing in $[-1, 0]$, in virtue of Proposition 5.1.

In this setting we prove the following: if $\mathfrak{h} \in \mathfrak{C}_{\mathbf{P}}^1$ is decreasing in $[-1, 0]$, no rotational \mathfrak{h} -tori exist. We follow the same argument exhibited by Tomter [Tom], pp. 490. Assume that $\alpha(s)$ generates a compact portion of an \mathfrak{h} -nodoid as in showed in Figure 3. Hence, $\alpha(s)$ is defined for $s \in [0, s_1]$, and satisfies $\theta(0) = \pi/2$, $\theta(s_0) = \pi$ and $\theta(s_1) = 3\pi/2$. Since \mathfrak{h} satisfies the property **(P)**, $\theta'(s) > 0$ for every $s \in [0, s_1]$. In this fashion, we express again the functions $x(s)$, $z(s)$ and $\nu(s)$ as functions of the angle θ , and the chain rule yields Equation (5.2).

Let us denote by $\theta \in [\pi/2, \pi]$ and $\alpha \in [\pi, 3\pi/2]$. Observe that both intervals are isometric under the orientation-reversing map $\alpha = \phi(\theta) = 2\pi - \theta$. In particular, if $\alpha = \phi(\theta)$, $\theta \neq \pi/2, \pi$, then $x(\theta) > x(\alpha)$ and $\nu(\theta) > \nu(\alpha)$, since $\nu(\theta) = \cos \theta / \sqrt{1 + \tau^2 x(\theta)^2}$. At this point, the chain rule yields

$$\frac{d(z \circ \phi)}{d\theta}(\theta) = -\frac{dz}{d\alpha}(2\pi - \theta) = \frac{-4 \sin(2\pi - \theta) \sqrt{1 + \tau^2 x(2\pi - \theta)^2}}{8\mathfrak{h}(\nu(2\pi - \theta)) - f(x(2\pi - \theta)) \sin(2\pi - \theta)}, \quad (5.4)$$

Since \mathfrak{h} is decreasing in $[-1, 0]$, for θ and $\alpha = \phi(\theta)$ as above we have $\mathfrak{h}(\nu(\theta)) \leq \mathfrak{h}(\nu(\alpha))$. Bearing these discussions in mind and Equation (5.4), the following inequality holds

$$\frac{d(z \circ \phi)}{d\theta}(\theta) = \frac{4 \sin(\theta) \sqrt{1 + \tau^2 x(2\pi - \theta)^2}}{8\mathfrak{h}(\nu(2\pi - \theta)) + f(x(2\pi - \theta)) \sin(\theta)} < \frac{4 \sin \theta \sqrt{1 + \tau^2 x(\theta)^2}}{8\mathfrak{h}(\nu(\theta)) - f(x(\theta)) \sin \theta} = \frac{dz}{d\theta}(\theta),$$

for every $\theta \in (\pi/2, \pi)$. In order to verify this inequality, take into account that

$$(f(x(\theta)) + f(x(2\pi - \theta))) \sin \theta > 8(\mathfrak{h}(\nu(\theta)) - \mathfrak{h}(\nu(2\pi - \theta))),$$

trivially holds since \mathfrak{h} is decreasing, and so the right-hand side is negative while the left-hand one is positive. Therefore, integrating for $\theta \in [\pi/2, \pi]$ we conclude

$$z(\pi) - z(3\pi/2) = \int_{\pi/2}^{\pi} \frac{d(z \circ \phi)}{d\theta}(\theta) d\theta < \int_{\pi/2}^{\pi} \frac{dz}{d\theta}(\theta) d\theta = z(\pi) - z(\pi/2).$$

This implies that $z(3\pi/2) > z(\pi/2)$ and no \mathfrak{h} -tori exist. Note that the particular choice $\mathfrak{h} = H_0 \in \mathbb{R}$ is compiled in this case, recovering the non-existence of rotational CMC tori in $\mathbb{H}^2(\kappa) \times \mathbb{R}$, Nil_3 and $\overline{SL}_2(\mathbb{R})$.

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The author was partially supported by P18-FR-4049.